



Status of XMASS



June 29th , 2011

Y. Kishimoto for XMASS collaboration

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- About XMASS project
- Current XMASS
 - Performance check
 - BG evaluation





ABOUT XMASS PROJECT

June 26, 2011

the 7th Patras workshop on
Axions, WIMPs and WISPs

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XMASS collaboration

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41 collaborators,
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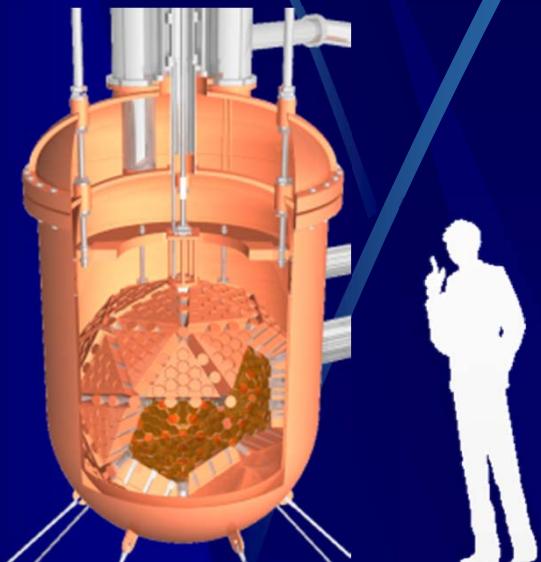
X MASS cite

- Kamioka mine -

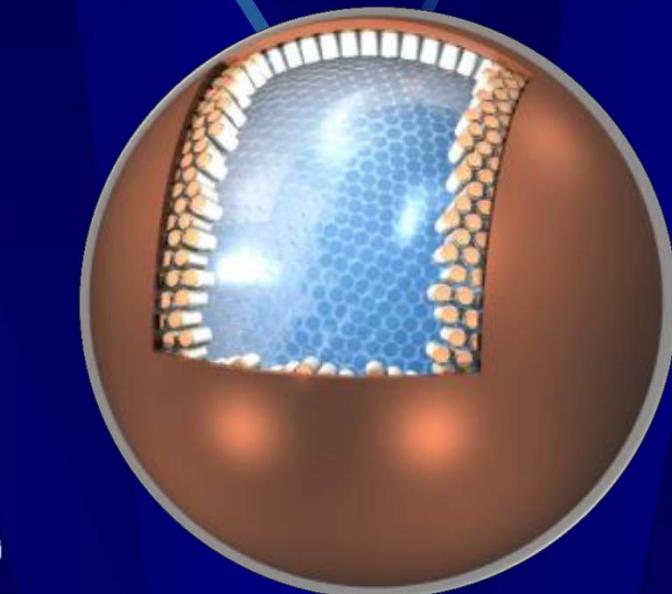


The XMASS experiment

- Xenon MASSive detector for solar neutrino (pp/ ^7Be)
- Xenon neutrino MASS detector (double beta decay)
- Xenon detector for Weakly Interacting MASSive particles (DM search)

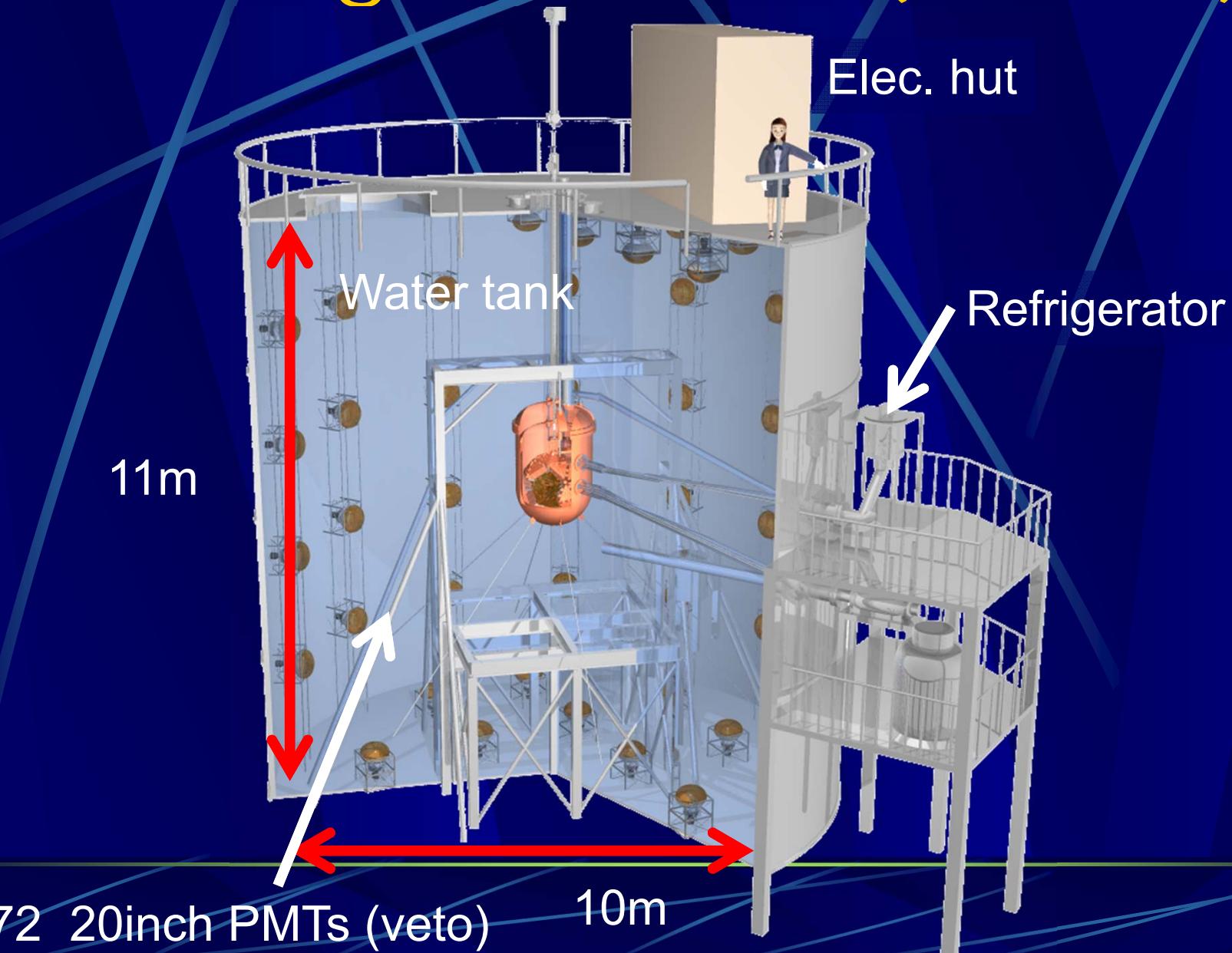


100kg FV (800kg LXe) ϕ 0.8m,
DM Current phase

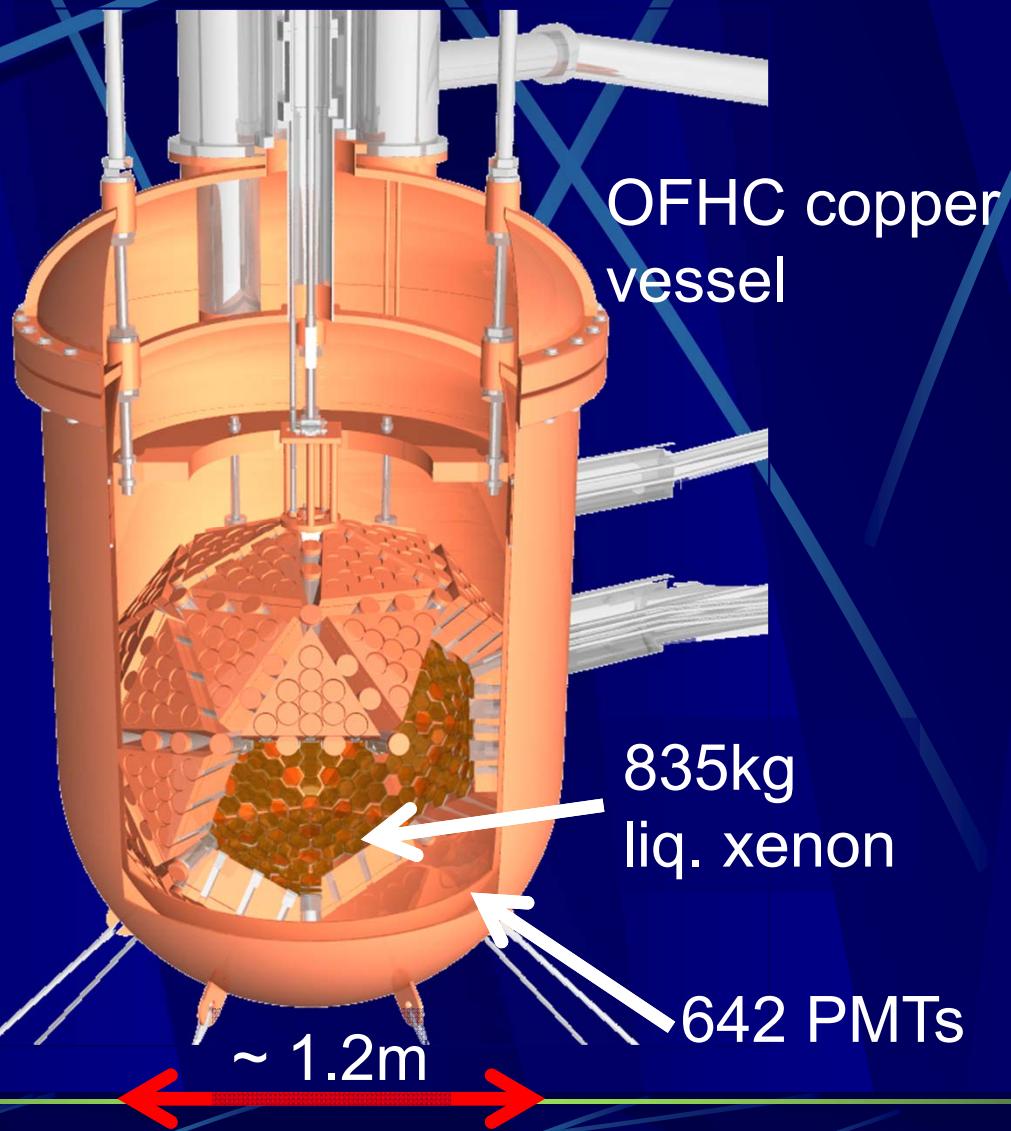


10ton FV (24ton) ϕ 2.5m
Solar ν , $0\nu\beta\beta$, DM in future

800kg detector (OD+ID)

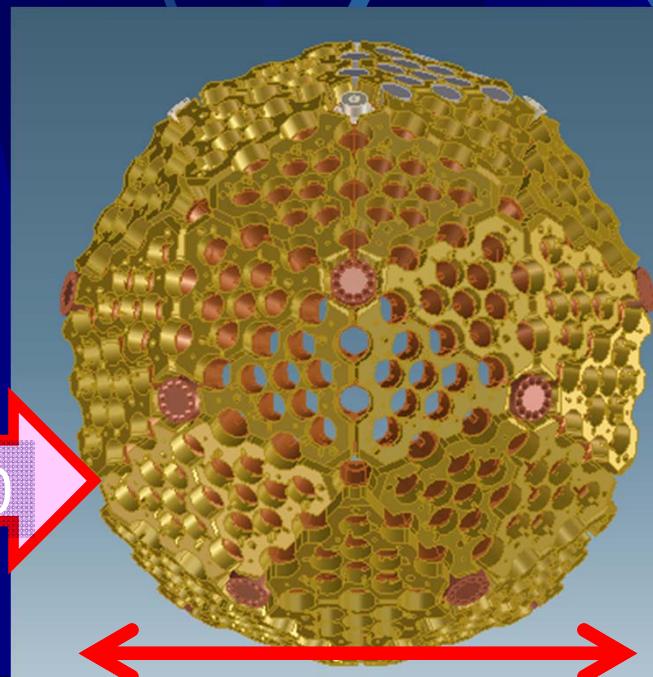
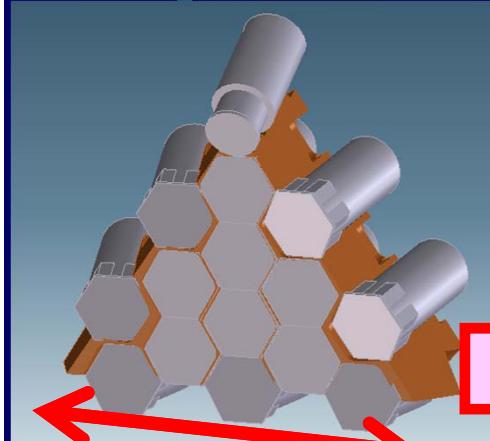


800kg detector (ID)



PMTs

- 642 PMTs :
 - 630 hex +12 round
- Q.E. : 28-39%
- Photo coverage: 62.4%



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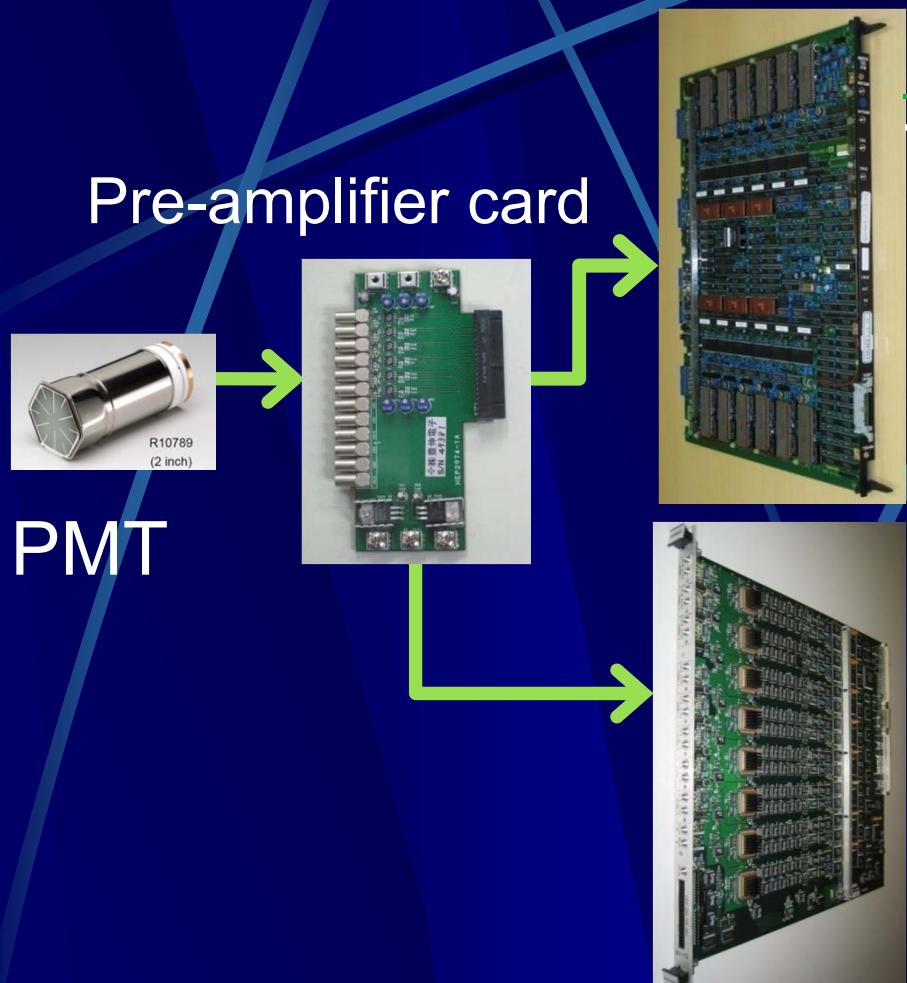
Hex: R10789-11



Round:
R10789-11MOD



DAQ hardware



12bit ADC/TDC (ATM)

To cover large energy range

- TKO module
- ADC dynamic range : 0~400 pC
- TDC dynamic range : 0~1 μ sec

8bit Flash ADC (60PMTs)

For pulse shape discrimination
in low energy

- dynamic range: 0~1 V
- sampling rate : 500 MS/s
- sample number: 8,160
- time span : 16.32 μ sec

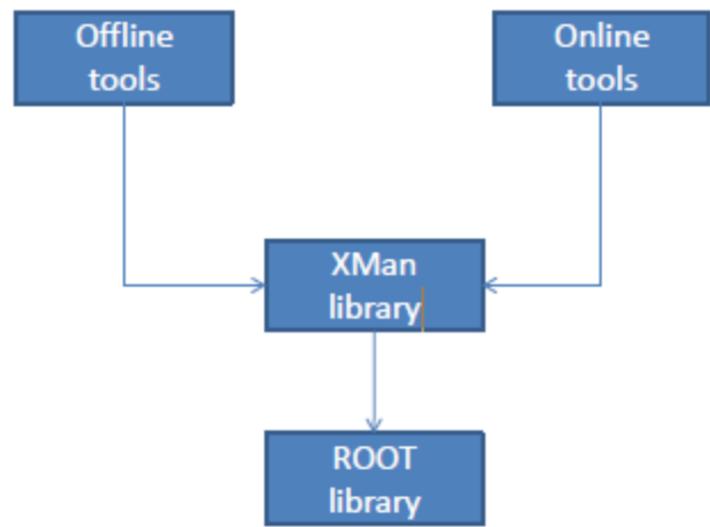
New FADCs for 642PMTs in November

10bit 1GS/s, 1V dynamic range

Analysis tool

- “XMAN”
 - Tools for XMASS based on ROOT
- Online and Offline analysis
- MC data analysis

XMASS analysis toolkit



BG reduction (I)

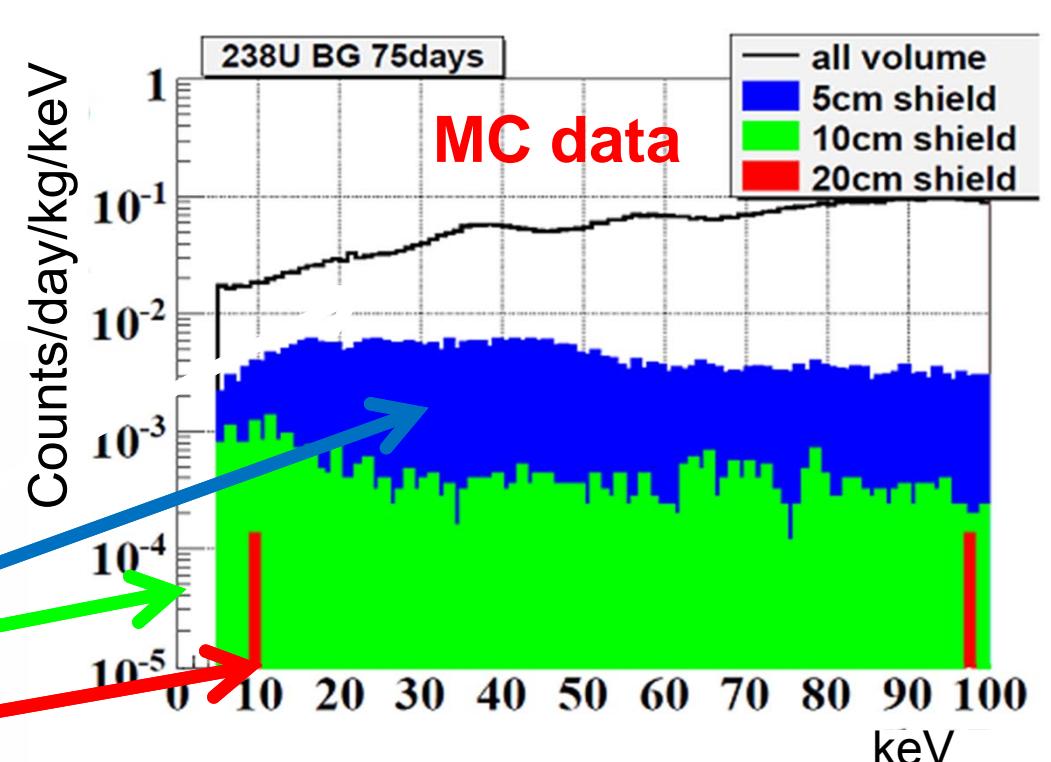
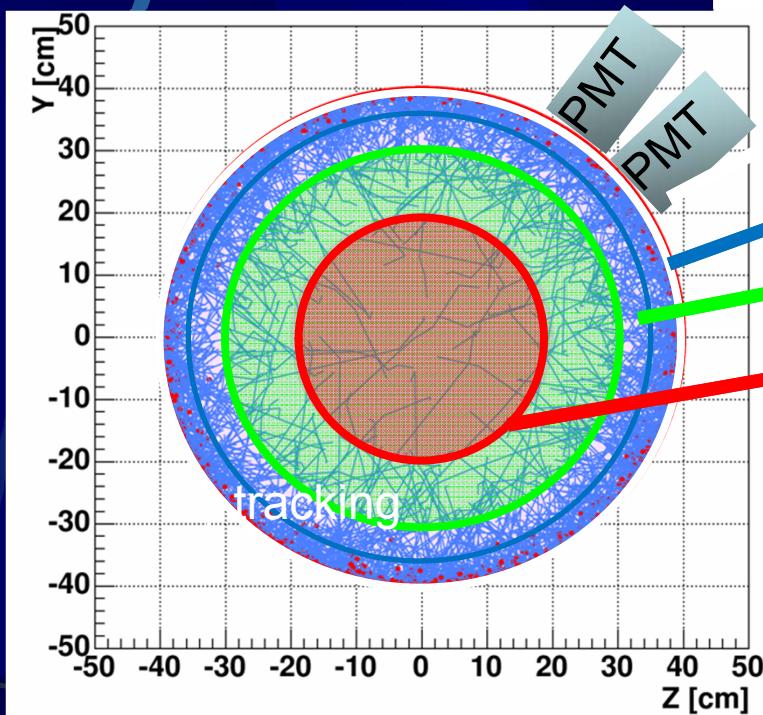
- material selection -

- PMTs
 - ~1/10 of familiar PMTs
- OFHC copper
 - They brought into the mine within one month after electro-refining.
- Other materials
 - All the components were selected with HPGe and ICP-MS.
 - >250 samples were measured.
 - →The total RI level is much lower than PMTs'.

BG reduction (II)

- self shielding -

| | BG/PMT [mBq] |
|------------------|---------------|
| U chain | 0.70 +/- 0.28 |
| Th chain | 1.51 +/- 0.31 |
| ⁴⁰ K | < 5.10 |
| ⁶⁰ Co | 2.92 +/- 0.16 |



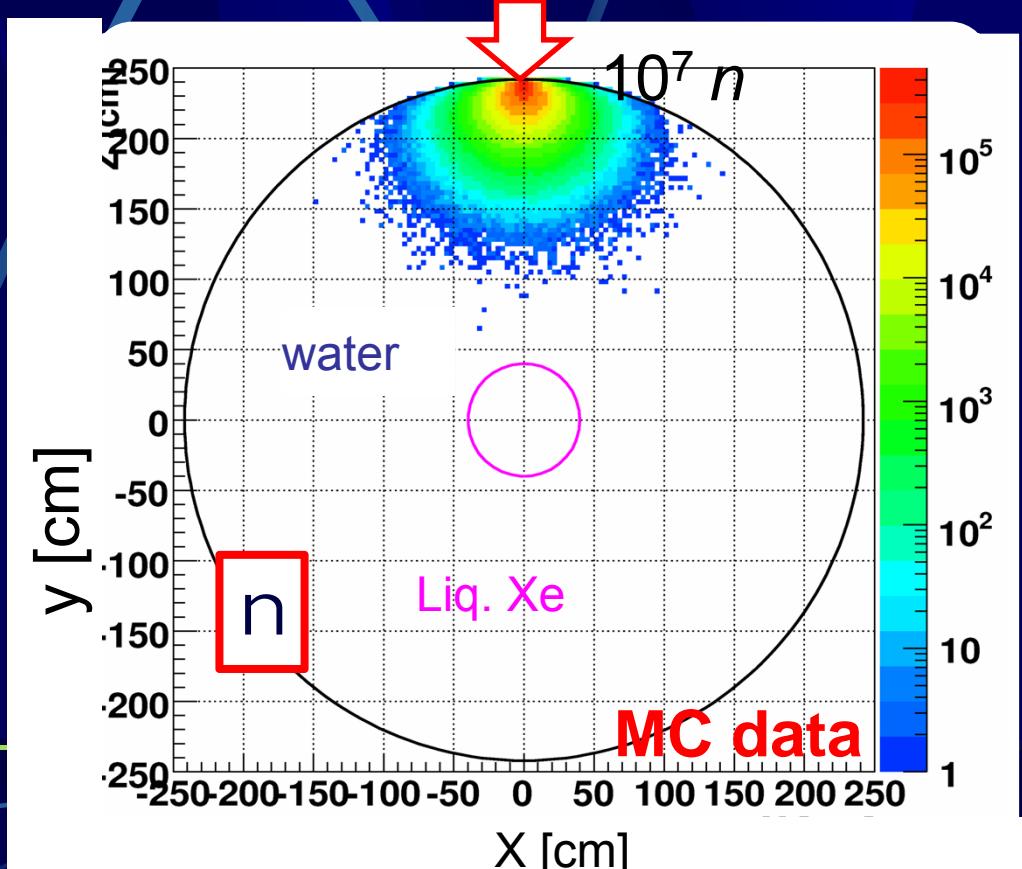
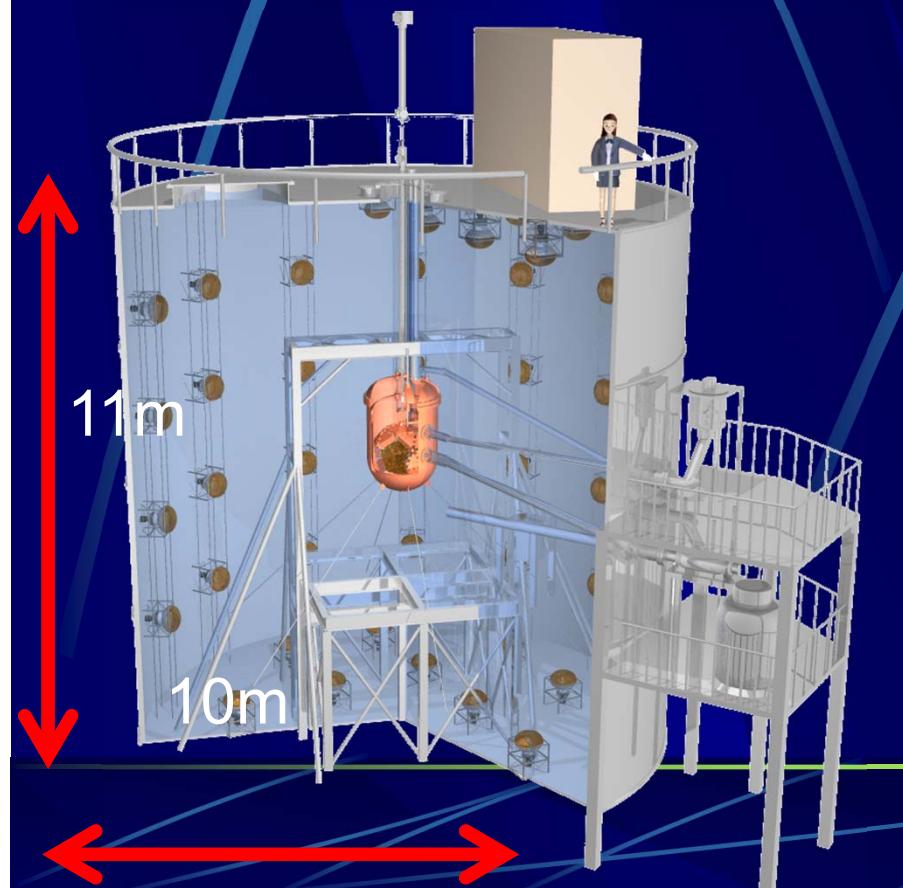
$< 10^{-4} / \text{keV/day/kg}$ (100kg F.V.)

n contribution $< 1.2 \times 10^{-5} / \text{d/kg/keV}$ (5-10keV)

BG reduction (III)

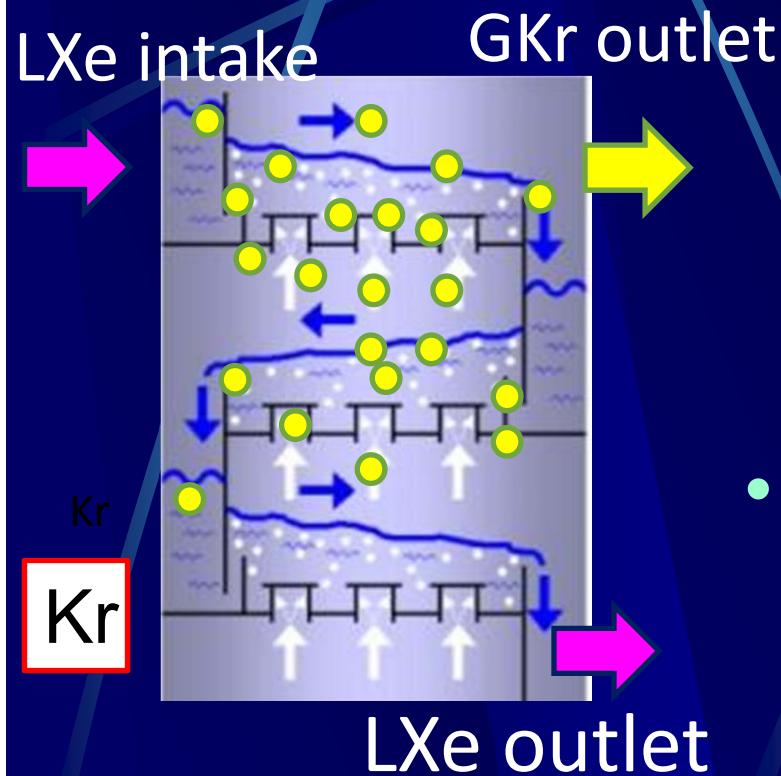
- Large Water Cherenkov OD -

- OD with 72 20"-PMTs \rightarrow active μ veto
- Large Vol. reduce γ and n



BG reduction(IV)

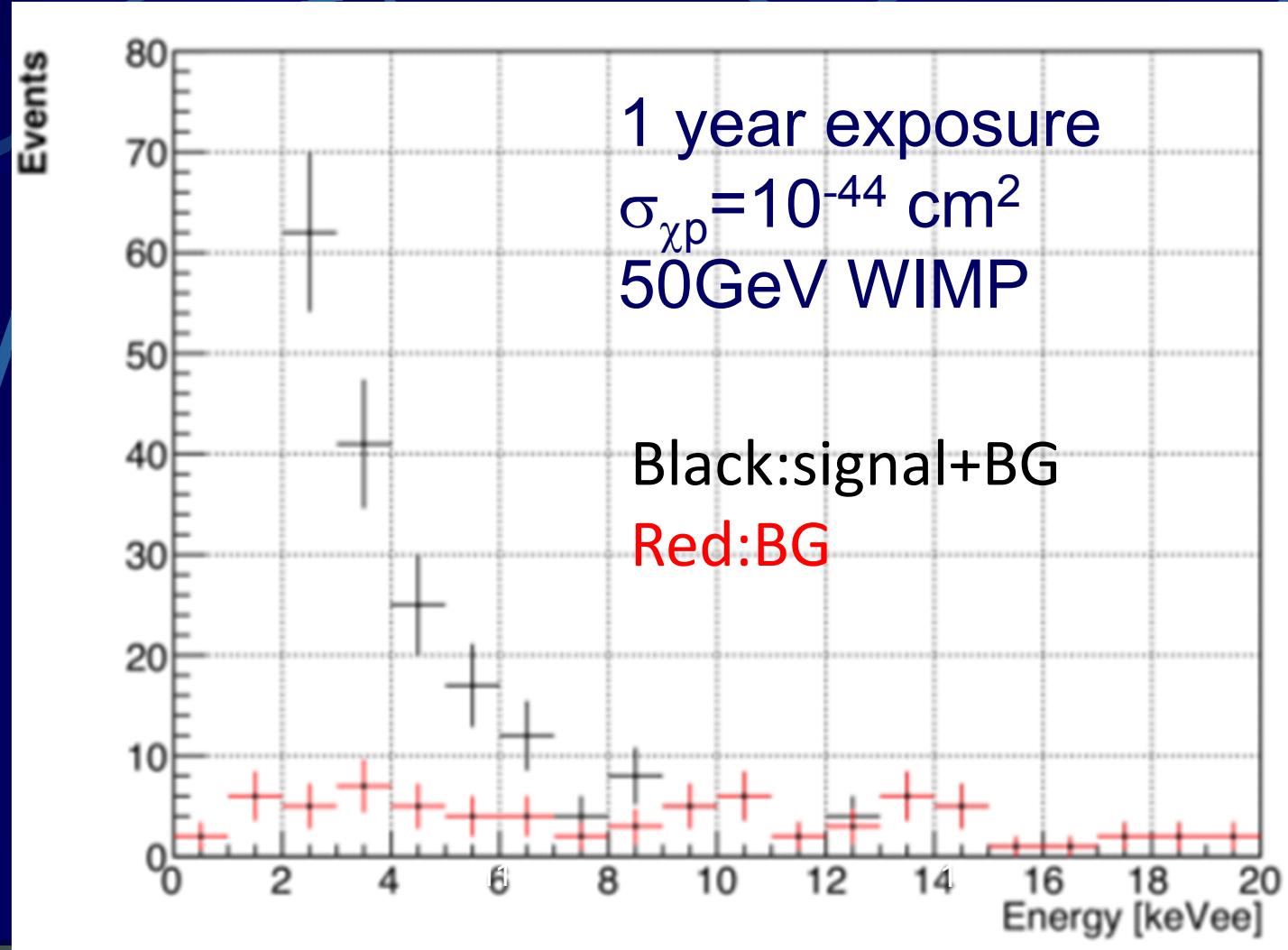
Xe purification



- Distillation:
 - 10^5 reduction for Kr (*)
 - $0.1\text{ppm} \rightarrow 1\text{ppt}$ for test sample
 - 1t of Xe was distilled at 10days before the filling.
- Other methods:
 - Getter
 - Filtering in liq. and gas phase is now under study for ^{222}Rn reduction.

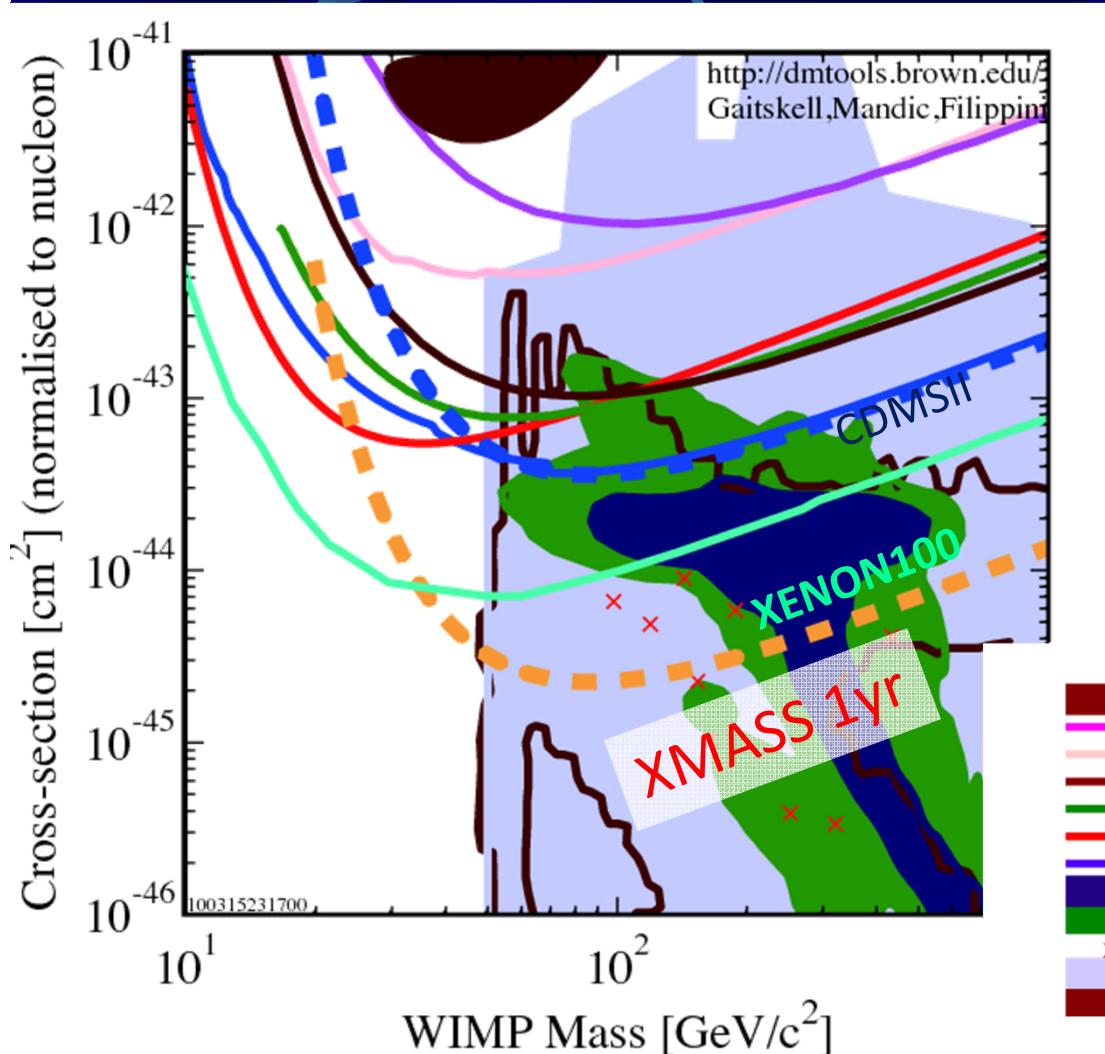
(*) K. Abe et al. for XMASS collab., Astropart. Phys. 31 (2009) 290

Expected spectrum



Expected sensitivity

Spin Independent



$\sigma_{\text{cp}} > 2 \times 10^{-45} \text{ cm}^2$
for 50-100GeV WIMP
(90% C.L.)

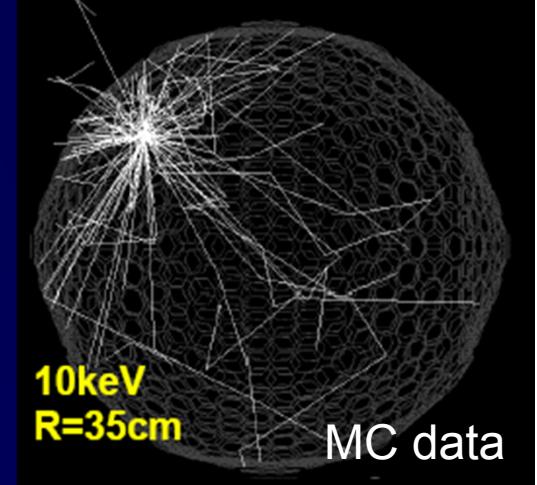
Assumption:

- ✓ 1yr exposure,
- ✓ 100kg FV
- ✓ BG: $1 \times 10^{-4} / \text{keV/day/kg}$
- ✓ Q factor: 0.2

DATA listed top to bottom on plot
DAMA/LIBRA 2008 3sigma, no ion channeling
WARP 2.3L, 96.5 kg-days 55 keV threshold
CRESST 2007 60 kg-day CaWO₄
Edelweiss II first result, 144 kg-days interleaved Ge
ZEPLIN III (Dec 2008) result
XENON10 2007, measured Leff from Xe cube
CDMS: Soudan 2004-2009 Ge
Trotta et al 2008, CMSSS Bayesian: 68% contour
Trotta et al 2008, CMSSS Bayesian: 95% contour
Ellis et. al Theory region post-LEP benchmark points
Baltz and Gondolo 2003
Baltz and Gondolo, 2004, Markov Chain Monte Carlos
100315231700

XMASS STATUS

- Performance check
- BG evaluation



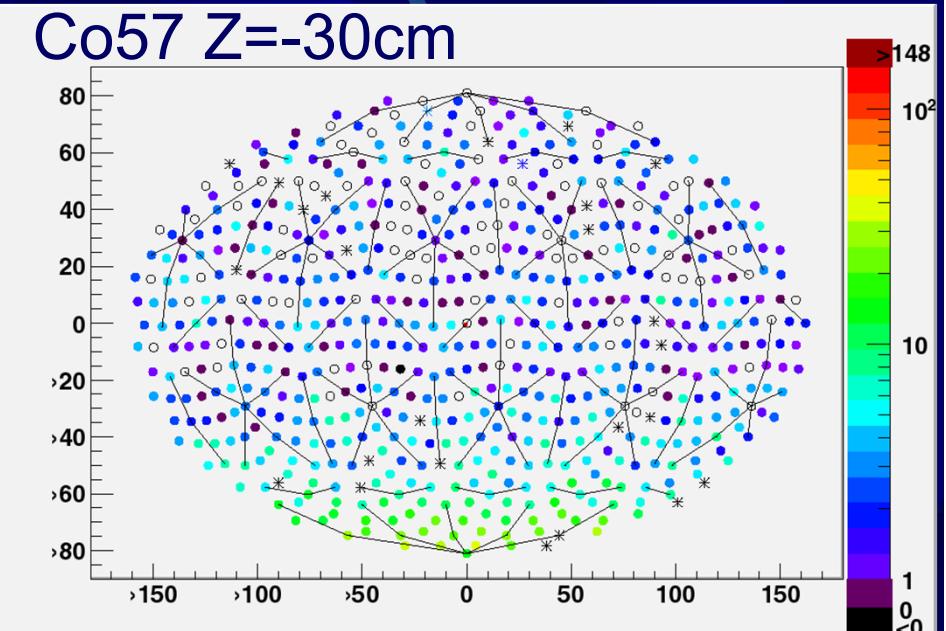
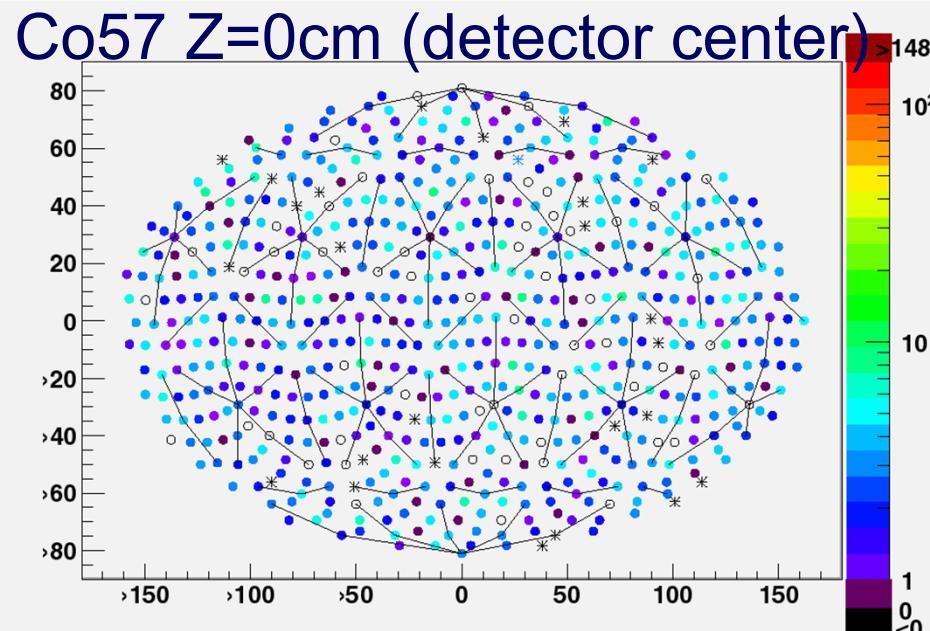
Construction of XMASS



- 2010 Feb.: PMT installation was finished.
- 2010 Sep.: Detector assembly was finished.
- 2010 Sep.: Distillation and liq. xenon filling.

Current status of XMASS

- Everything was ready.
 - Detector itself
 - DAQ hardware and software
 - Calibration apparatus and so on
- 2010 Oct. → **Commissioning phase**



Calibration system

- Features
 - <1mm precision
 - Reducing “shadowing” by thin wire source
- Sources
 - ^{57}Co , ^{241}Am , ^{109}Cd , ^{55}Fe , ^{137}Cs , ...

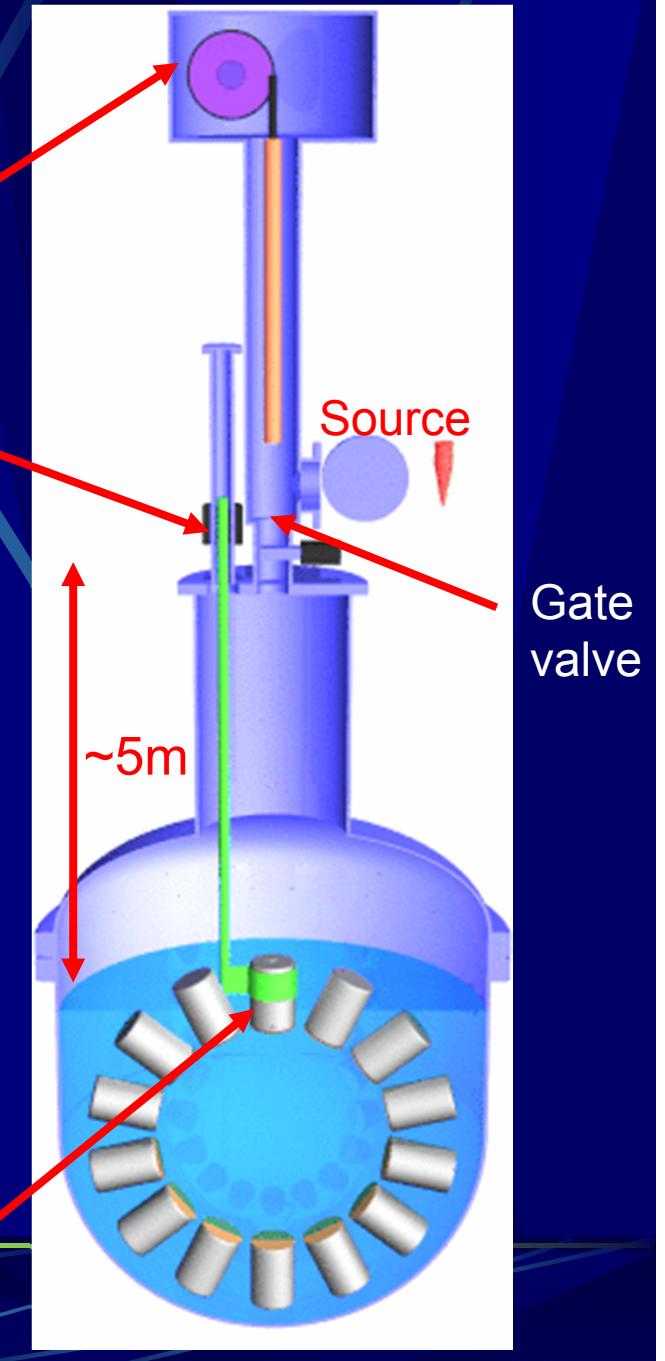
0.21mm ϕ for ^{57}Co source



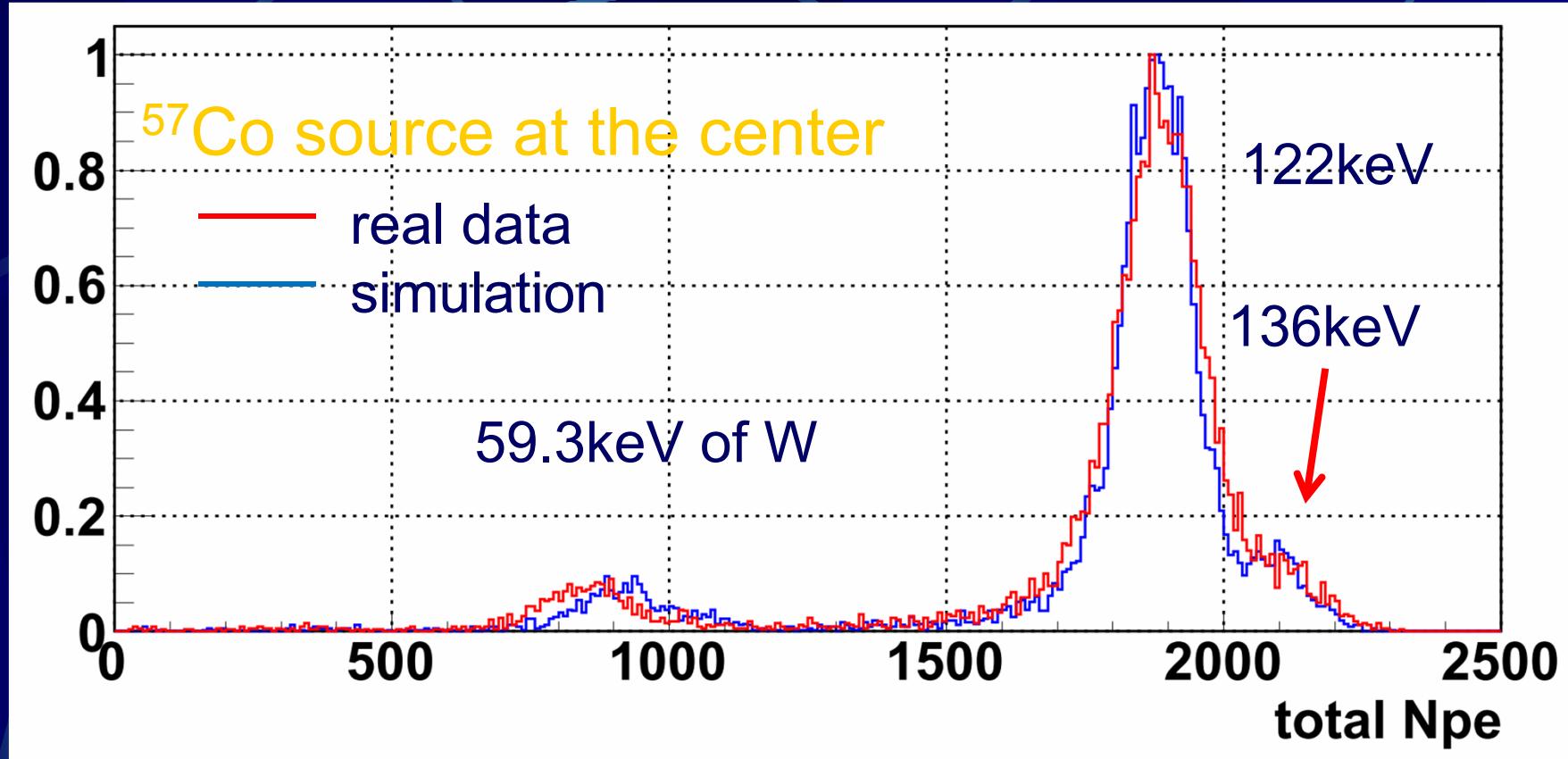
Source rod with a dummy source

Stepping Motor
Linear Motion Feed-through

Top photo tube



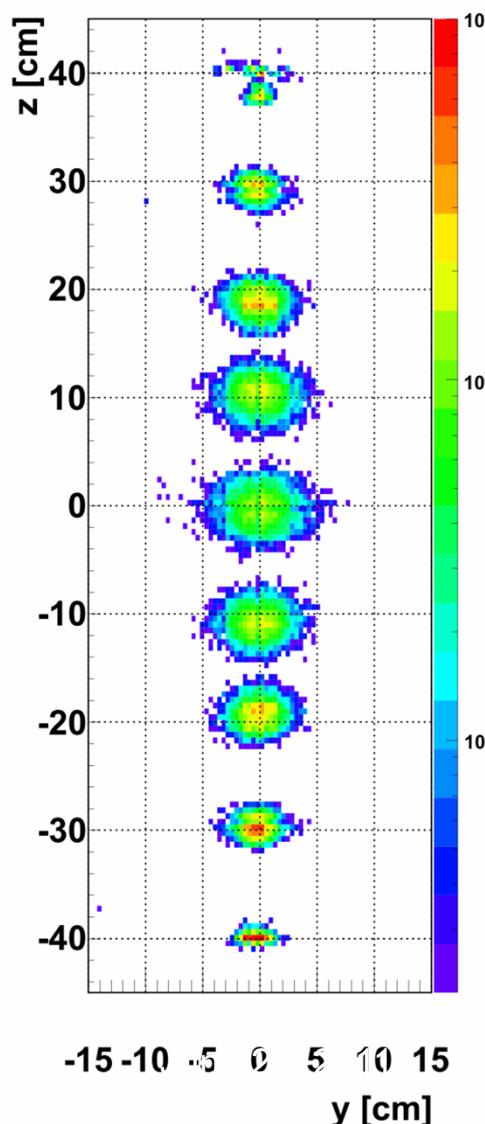
PE distribution



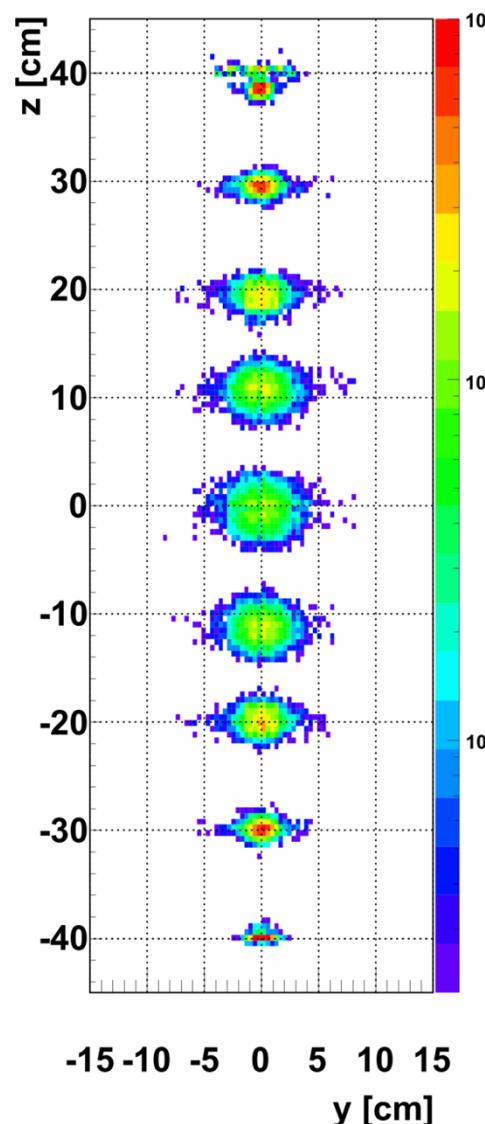
- High p.e. yield, 15.1 ± 1.2 p.e./keV, was obtained.
- The photo electron yield distribution was well reproduced by MC.

Vertex reconstruction

Real Data



Simulation



r is well reconstructed.

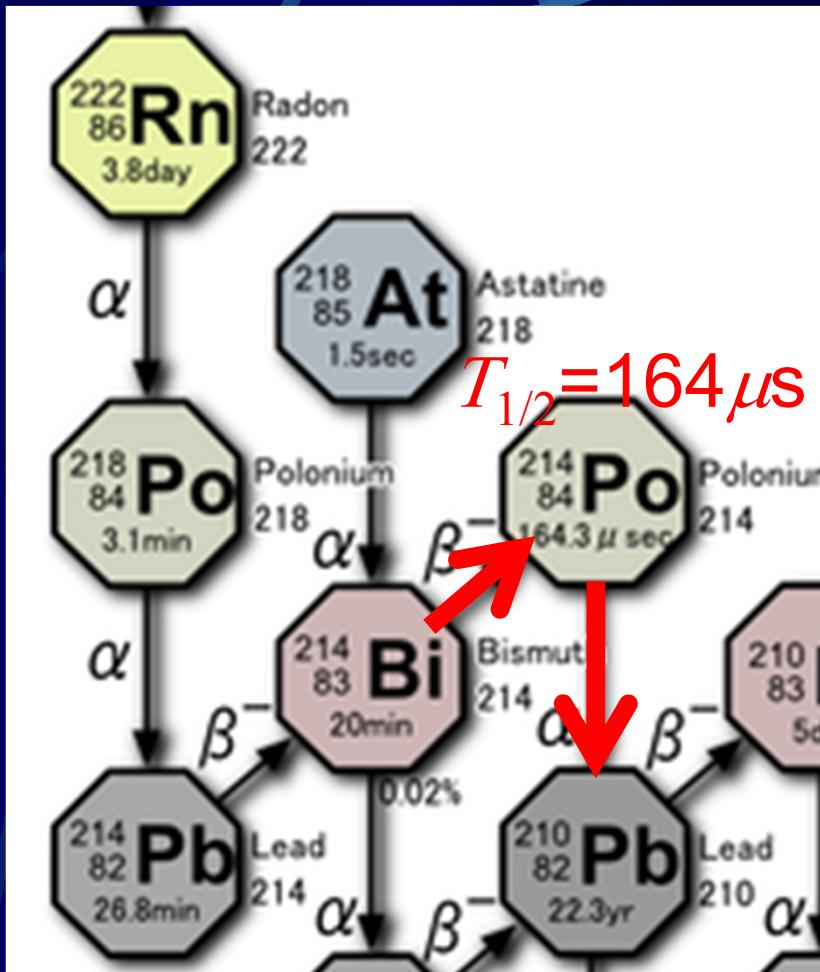
Δr is as expected by MC
1.4cm RMS @ $z=0\text{cm}$
1.0cm RMS @ $\pm 20\text{cm}$

For 122keV γ rays

Evaluation of internal BG

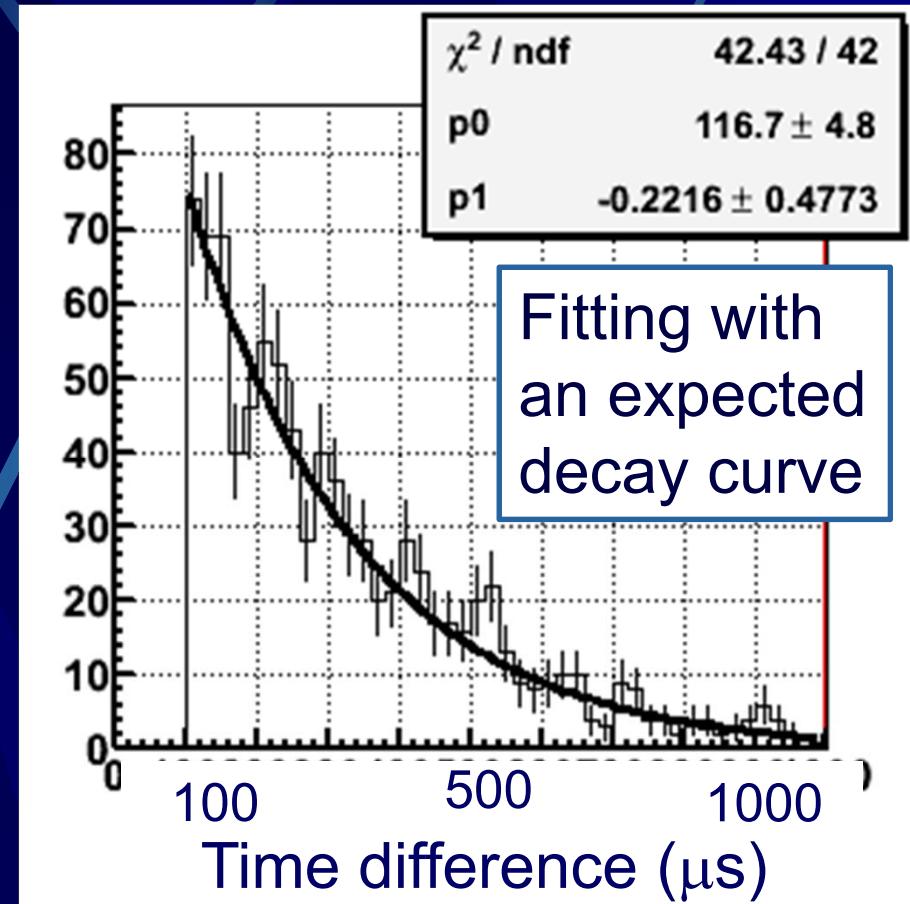
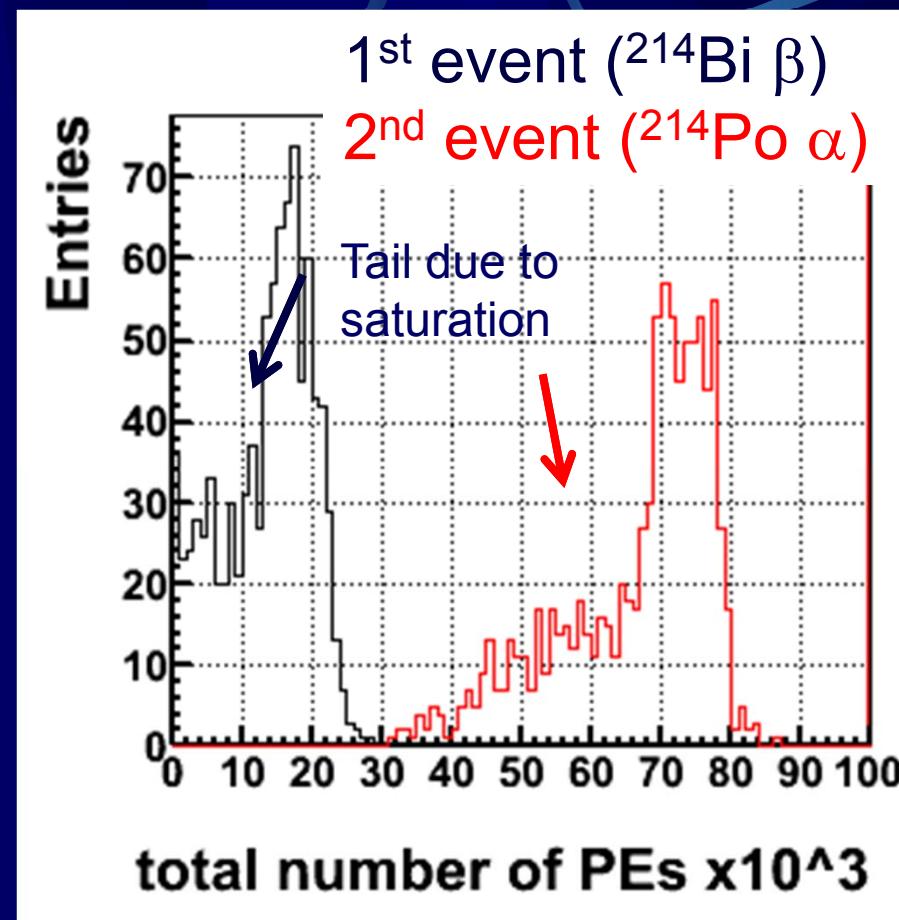
- External BG:
 - External BG (γ, n) can be effectively reduced by the water tank and the outer part of LXe.
- Internal BG:
 - Internal BG must be reduced by other means.
 - ^{222}Rn , ^{220}Rn , and ^{85}Kr are our concern because they give low energy BG.

^{222}Rn measurement



- Delayed coincidence
- $T_{1/2} = 164 \mu\text{s}$
- The 2nd α is very bright, so gain of 321 PMTs are reduced for wide dynamic range.

222Rn Result



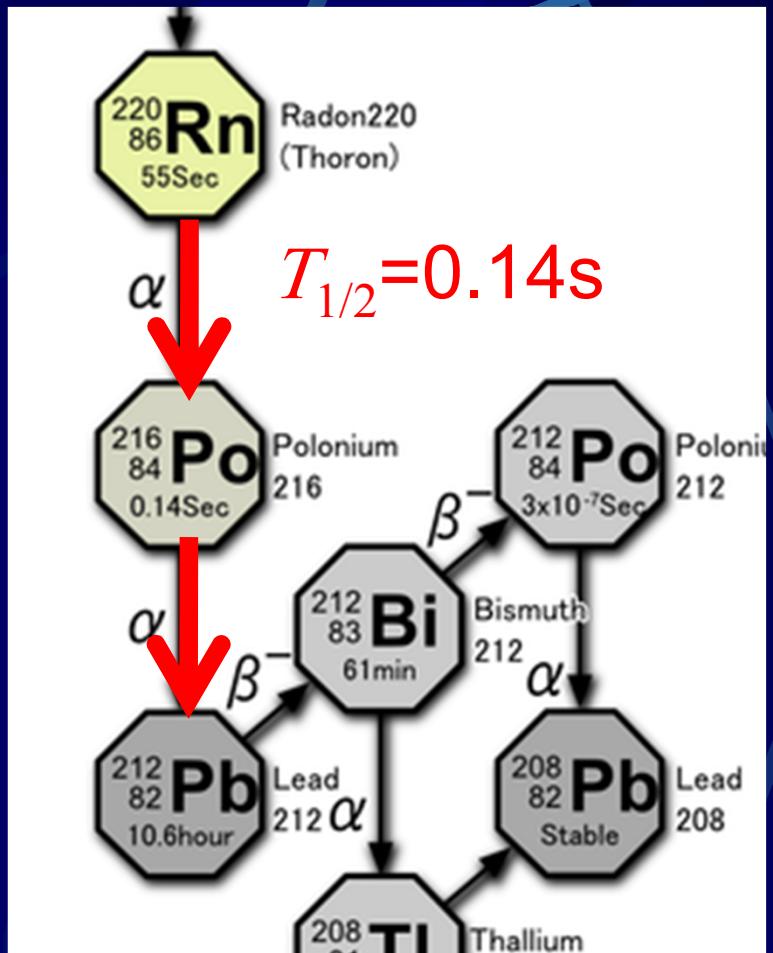
$^{222}\text{Rn} = 8.2(0.5)\text{mBq}$ in inner volume

Target value = 1 mBq.

→ Investigating event distribution

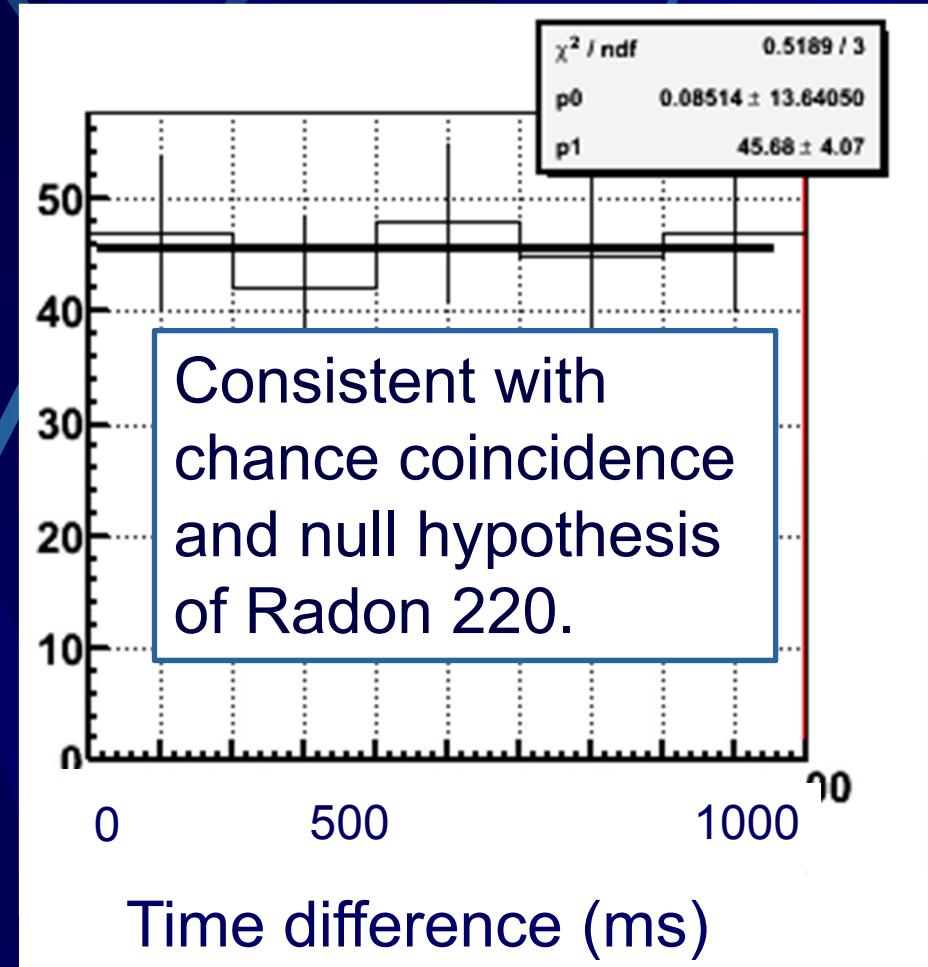
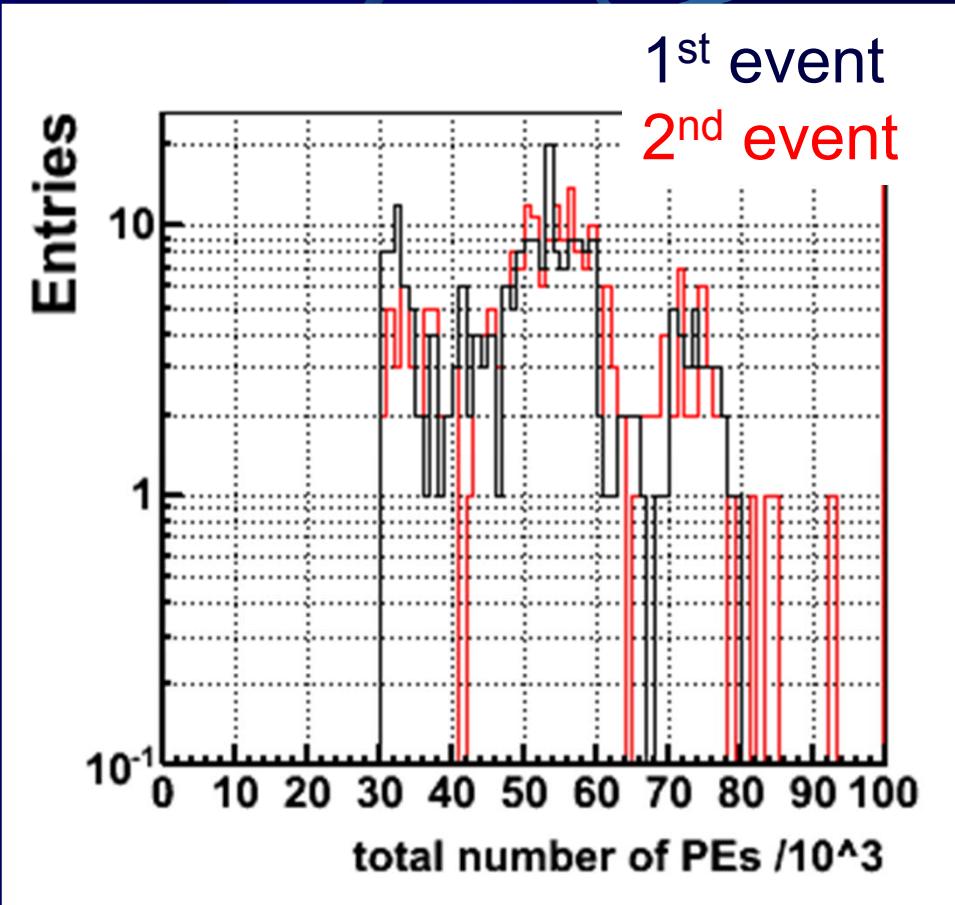
Considering measures to remove (Charcoal Trap)

^{220}Rn measurement



- Delayed coincidence
- $T_{1/2}=0.14\text{s}$

^{220}Rn Result

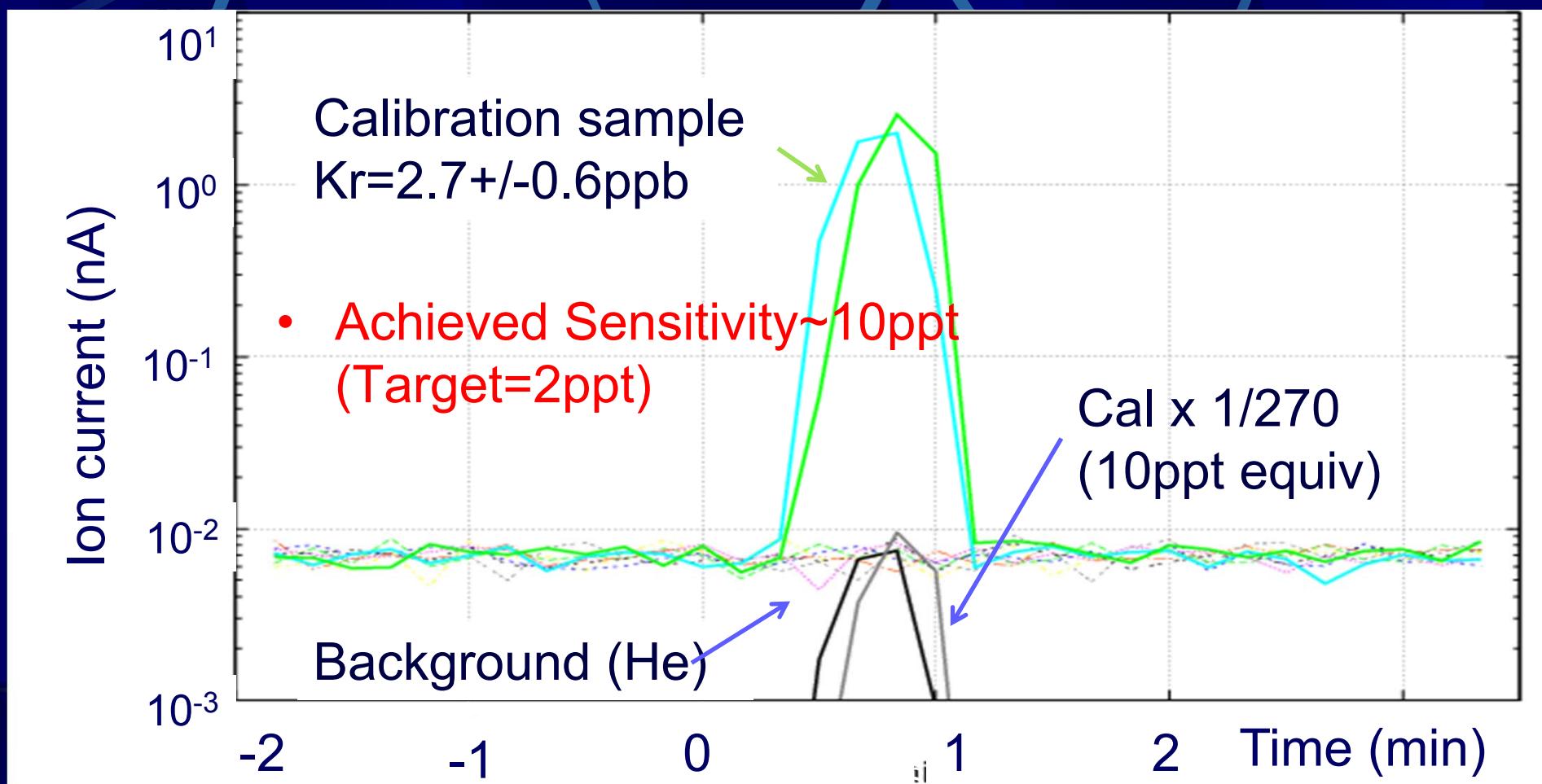


- $^{220}\text{Rn} < 0.28\text{mBq}$ in inner volume

Lower than our target value = 0.43mBq .

Kr measurement

- Gas Chromatography + Atmospheric Pressure Ionization Mass Spectrometer
- ^{85}Kr can be estimated with measured $^{85}\text{Kr}/\text{Kr}$

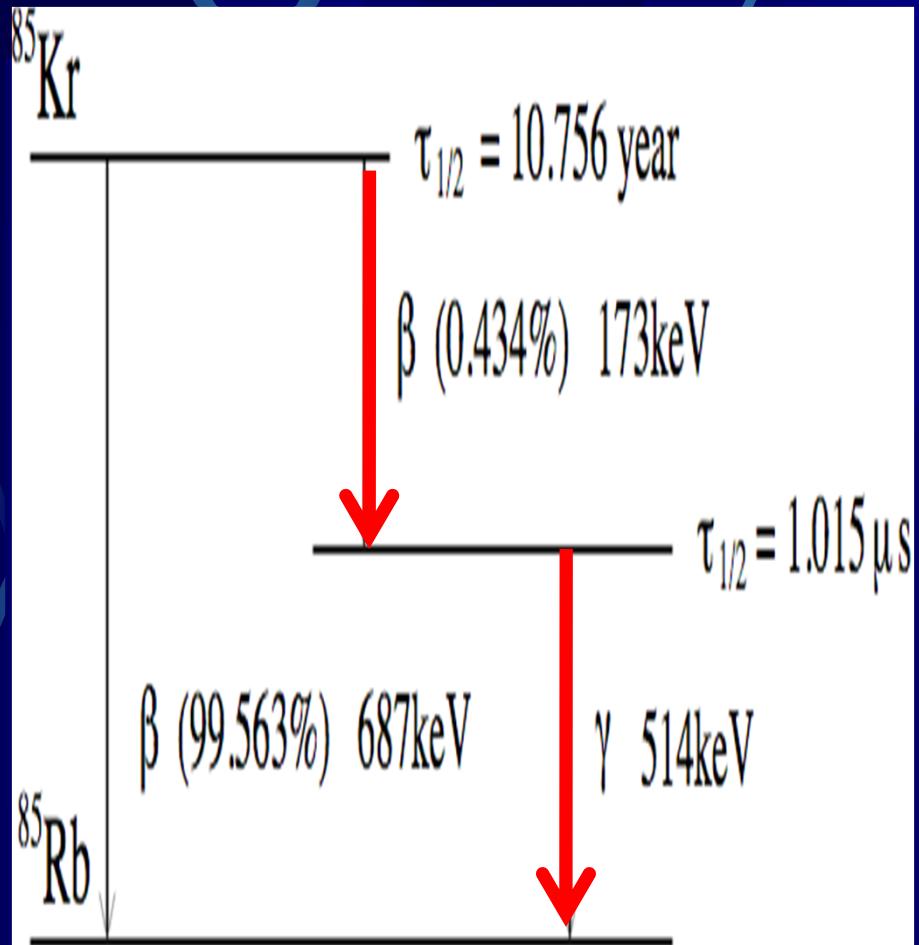


Kr measurement (II)

- Delayed coincident
- $T_{1/2} = 1.0 \mu\text{s} \rightarrow$

^{85}Kr measurement
with **FADC**

- New 650 FADCs in November
 - They will boost our detailed studies.



Summary

- The XMASS 800kg detector is for DM search with the sensitivity $2 \times 10^{-45} \text{ cm}^2$ (spin independent case).
- Construction of the 800kg detector finished.
- Commissioning runs are on going to confirm performance and BG properties.
 - Energy resolution and vertex resolution were as expected.
 - $\sim 1\text{cm}$ position resolution and $\sim 4\%$ energy resolution for $122\text{keV} \gamma$.
 - Radon background are close to the target values
 - Kr contamination will be evaluated soon.
- Results are coming after commissioning.

BACKUP SLIDES

June 26, 2011

the 7th Patras workshop on
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Vertex reconstruction

- $L(\mathbf{r}) = \prod_{i \in \text{PMT}} \exp(-\mu(i)) \mu(i)^{n(i)} / n(i)!$
 - $\mu(i) = F(\mathbf{r}, i) / \sum_{i \in \text{PMT}} F(\mathbf{r}, i) \times \text{total NPE}$
 - $n(i) = \text{observed NPE in } i\text{-th PMT}$
 - $F(\mathbf{r}, i)$: Acceptance of i -th PMT
 - This is calculated by MC simulation.
- Find \mathbf{r} which gives min. $\log L(\mathbf{r})$

Energy reconstruction

- Calculated from
 - observed NPE
 - and total acceptance $\sum_{i=\text{PMT}} F(r, i)$

Why LXe?

- High Atomic mass Xe ($A \sim 131$) good for SI case ($\sigma \propto A^2$).
- Odd isotope (^{129}Xe (26.4%), ^{131}Xe (21.3%)) with large SD enhancement factors.
- High atomic number ($Z=54$) and density ($\rho = \sim 3\text{g/cc}$)
 - -> Effective self-shielding.
 - -> Compact for large mass detector.
- High photon yield
- Easy to purify for both electro-negative and radioactive impurities.
 - -> By circulation of Xe with getter for electro-negative.
 - -> Distillation for ^{85}Kr removal.

Reconstructed energy

^{57}Co source at the center

