



Status of XMASS

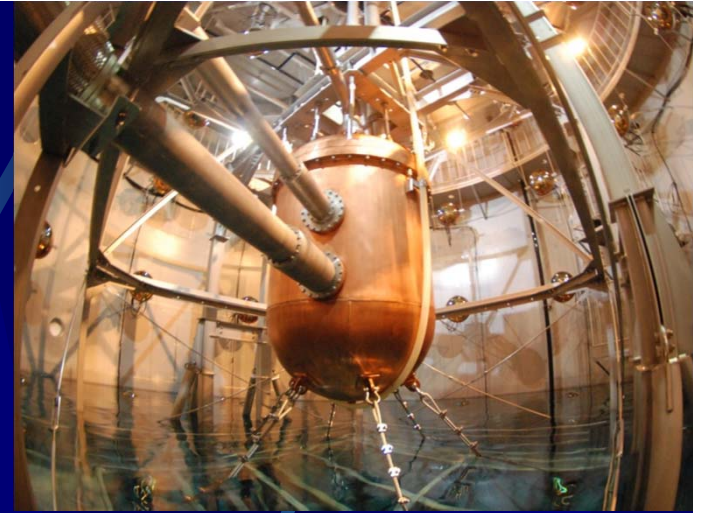
June 29th , 2011

Y. Kishimoto for XMASS collaboration

Contents

- About XMASS project
- Current XMASS
 - Performance check
 - BG evaluation





ABOUT XMASS PROJECT

June 26, 2011

the 7th Patras workshop on
Axions, WIMPs and WISPs

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XMASS collaboration

Kamioka Observatory, ICRR, Univ. of Tokyo :

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K. Hiraide, A. Shinozaki, S. Hirano, D. Umemoto, O. Takachio, K. Hieda

IPMU, University of Tokyo : K. Martens, J.Liu

Kobe University: Y. Takeuchi, K. Otsuka, K. Hosokawa, A. Murata

Tokai University: K. Nishijima, D. Motoki, F. Kusaba

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**41 collaborators,
10 institutes**

Gifu pre.
Hida-city
Kamioka-cho
Ikenoyama Mt.

X MASS cite

- Kamioka mine -

Candles
Kamland
XMASS
SG
CLIO
Super
Kamiokande

1000m

(2700m w.e.)

~360m

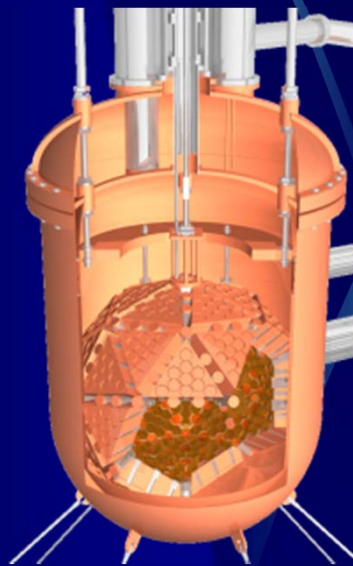
Atotsu Entrance

To Tateyama
Atotsu River

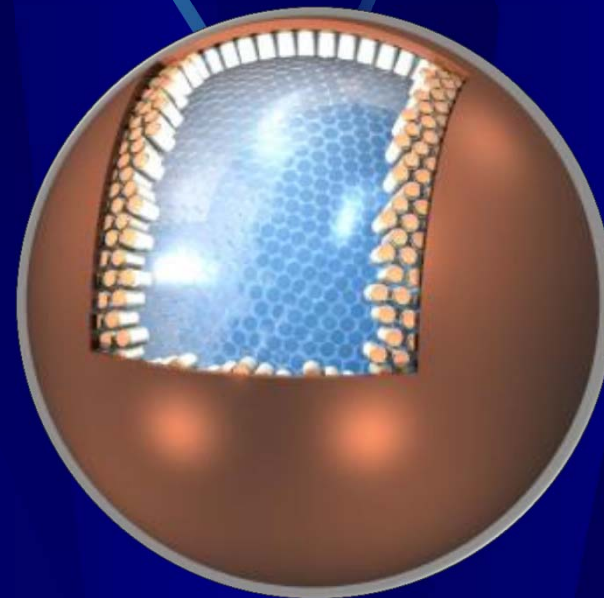


The XMASS experiment

- Xenon **MASS**ive detector for solar neutrino ($pp/{}^7\text{Be}$)
- Xenon neutrino **MASS** detector (double beta decay)
- Xenon detector for Weakly Interacting **MASS**ive particles (DM search)

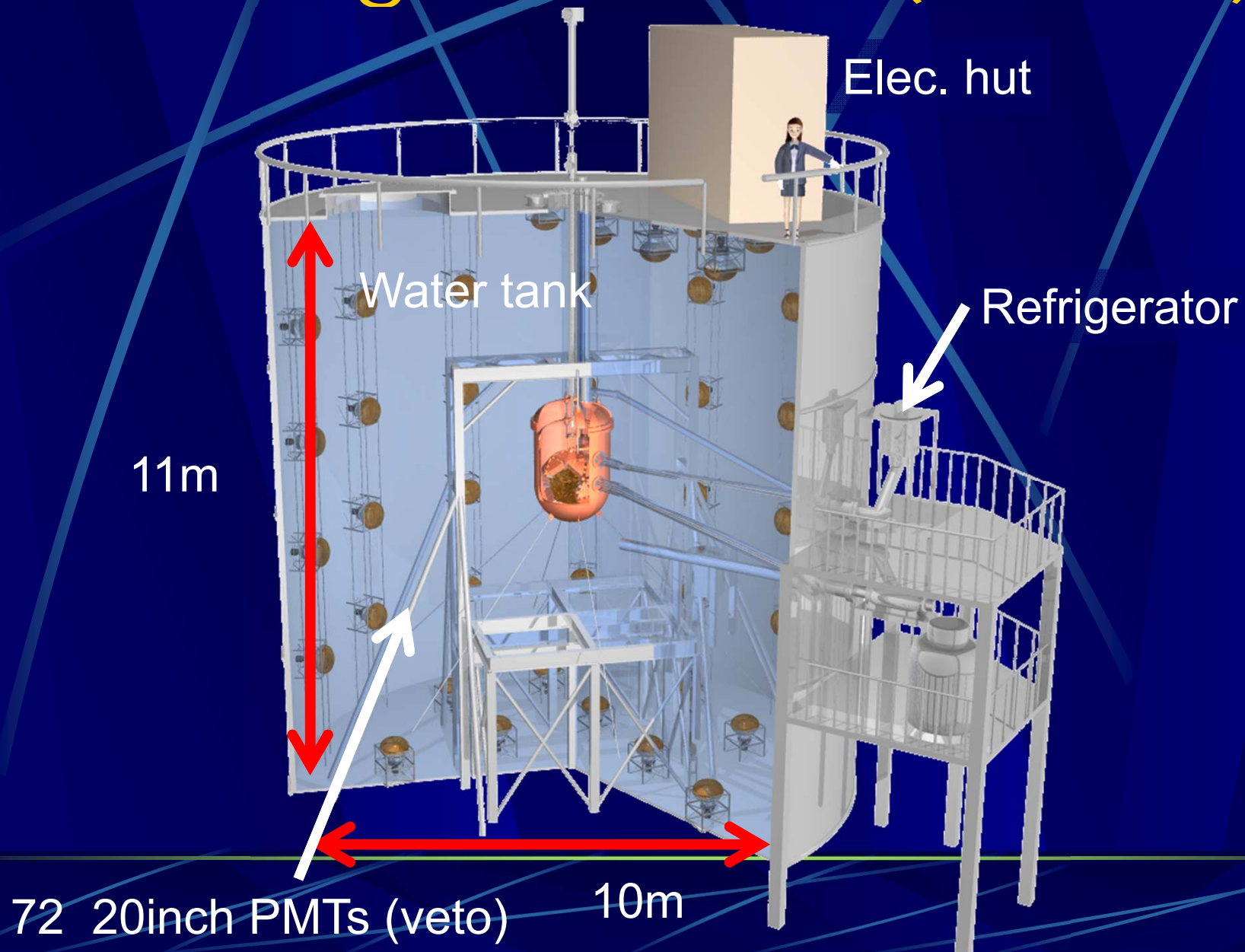


100kg FV (800kg LXe) ϕ 0.8m,
DM Current phase

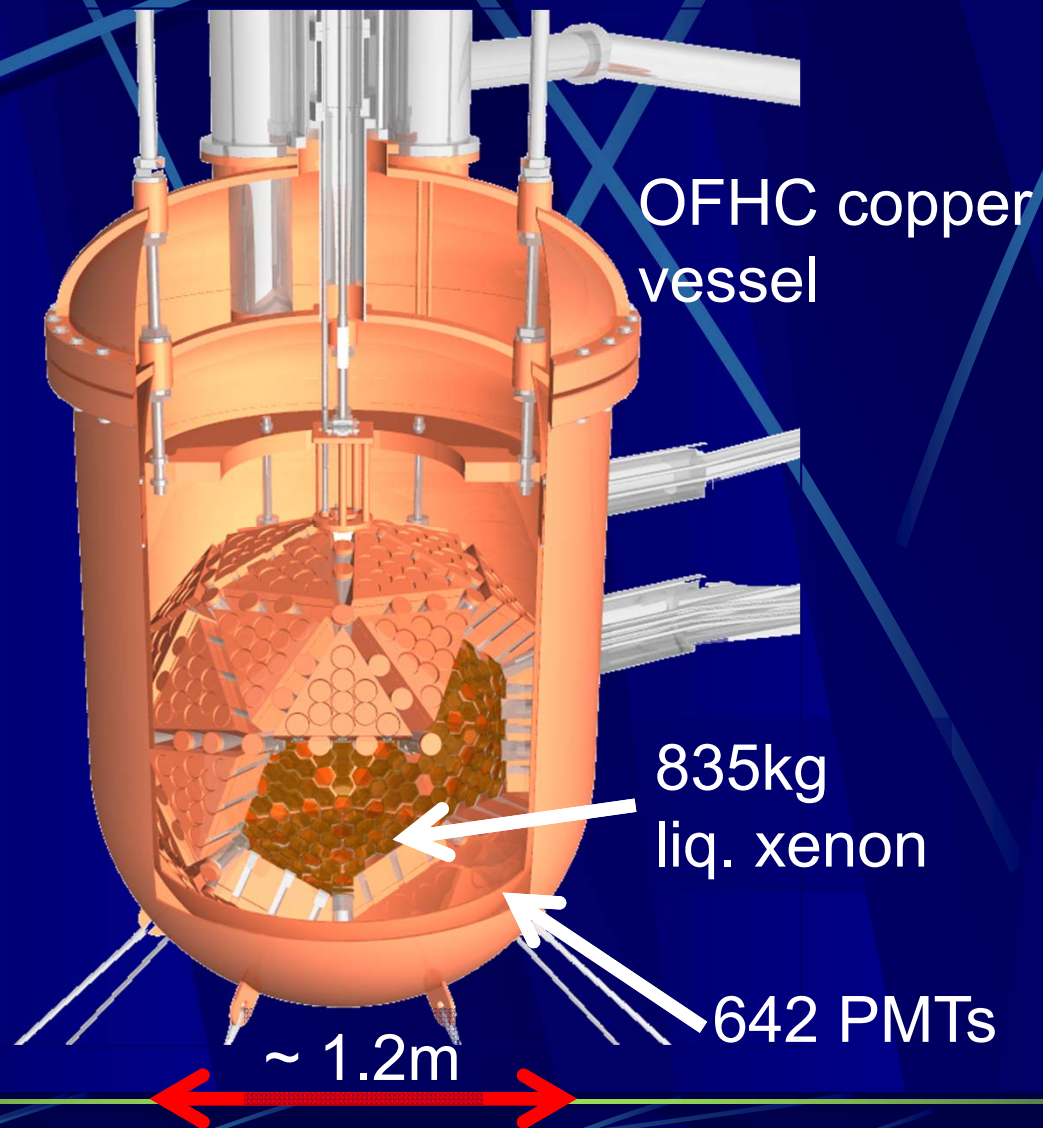


10ton FV (24ton) ϕ 2.5m
Solar ν , $0\nu\beta\beta$, DM in future

800kg detector (OD+ID)

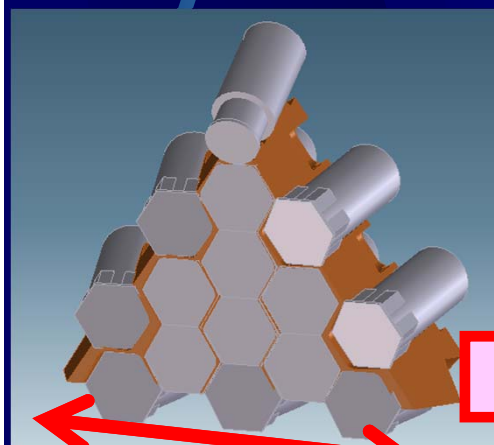


800kg detector (ID)



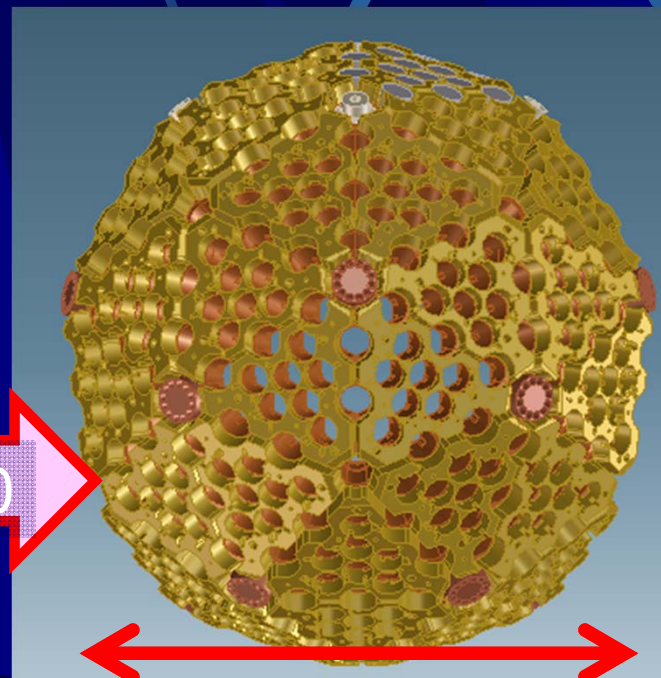
PMTs

- 642 PMTs :
 - 630 hex + 12 round
- Q.E. : 28-39%
- Photo coverage: 62.4%



310mm

×60



$\Phi 1113\text{mm}$

Hex: R10789-11



Round:
R10789-11MOD



June 26, 2011

DAQ hardware



PMT

Pre-amplifier card



12bit ADC/TDC (ATM)

To cover large energy range

- TKO module
- ADC dynamic range : 0~400 pC
- TDC dynamic range : 0~1 μ sec

8bit Flash ADC (60PMTs)

For pulse shape discrimination in low energy

- dynamic range: 0~-1 V
- sampling rate : 500 MS/s
- sample number: 8,160
- time span : 16.32 μ sec

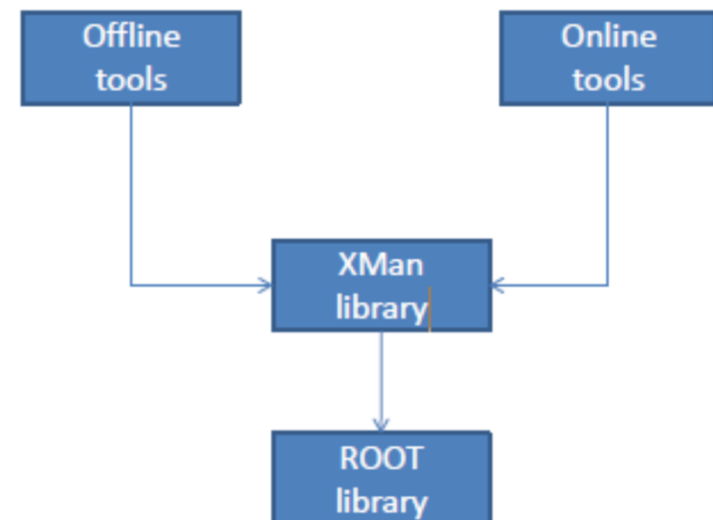
New FADCs for 642PMTs in November

10bit 1GS/s, 1V dynamic range

Analysis tool

- “XMAN”
 - Tools for XMASS based on ROOT
- Online and Offline analysis
- MC data analysis

XMASS analysis toolkit



BG reduction (I)

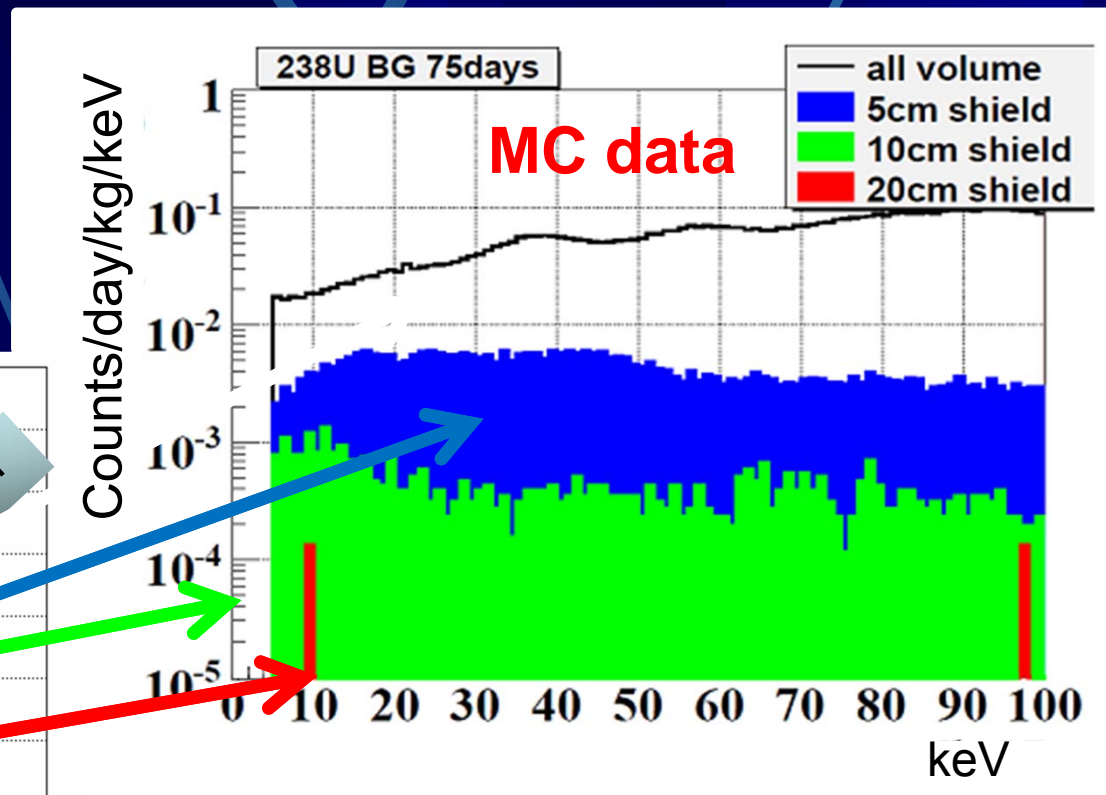
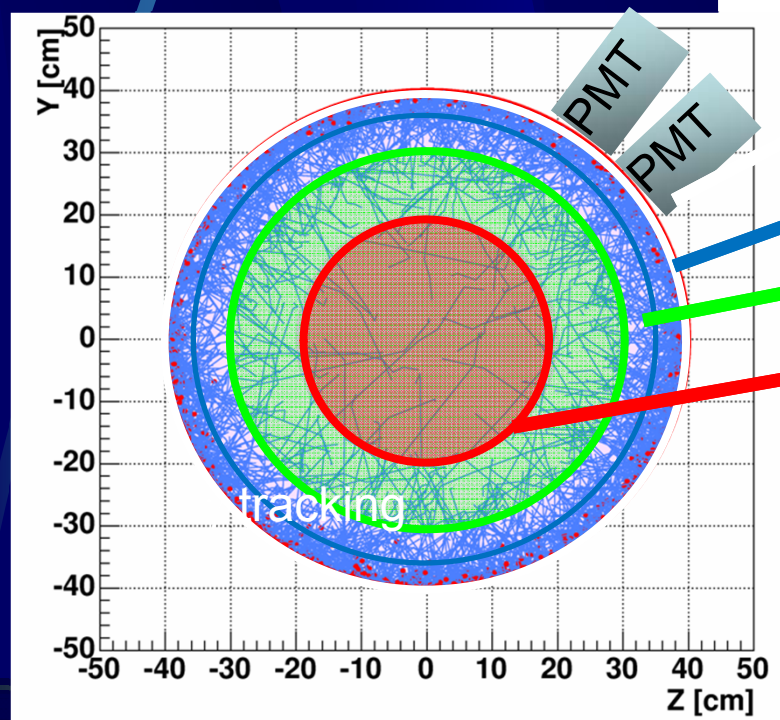
- material selection -

- PMTs
 - ~1/10 of familiar PMTs
- OFHC copper
 - They brought into the mine within one month after electro-refining.
- Other materials
 - All the components were selected with HPGe and ICP-MS.
 - >250 samples were measured.
 - → The total RI level is much lower than PMTs'.

BG reduction (II)

- self shielding -

	BG/PMT [mBq]
U chain	0.70 +/- 0.28
Th chain	1.51 +/- 0.31
^{40}K	< 5.10
^{60}Co	2.92 +/- 0.16

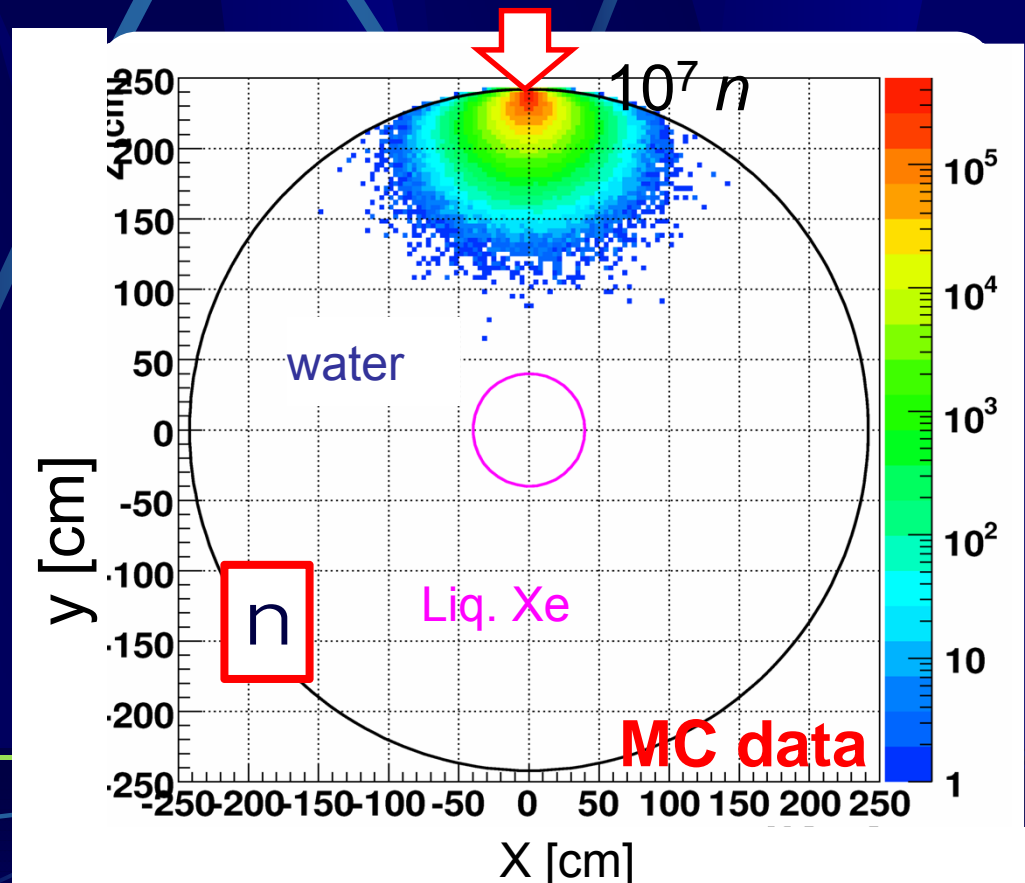
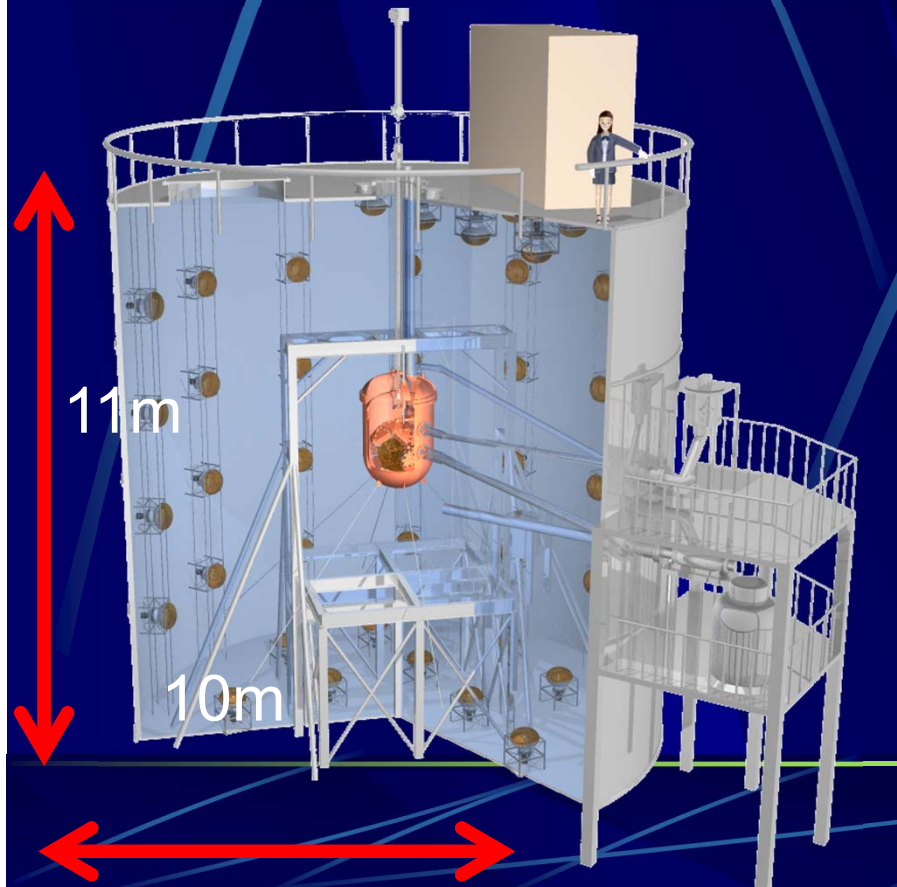


< 10^{-4} /keV/day/kg (100kg F.V.)
n contribution < 1.2×10^{-5} /d/kg/keV (5-10keV)

BG reduction (III)

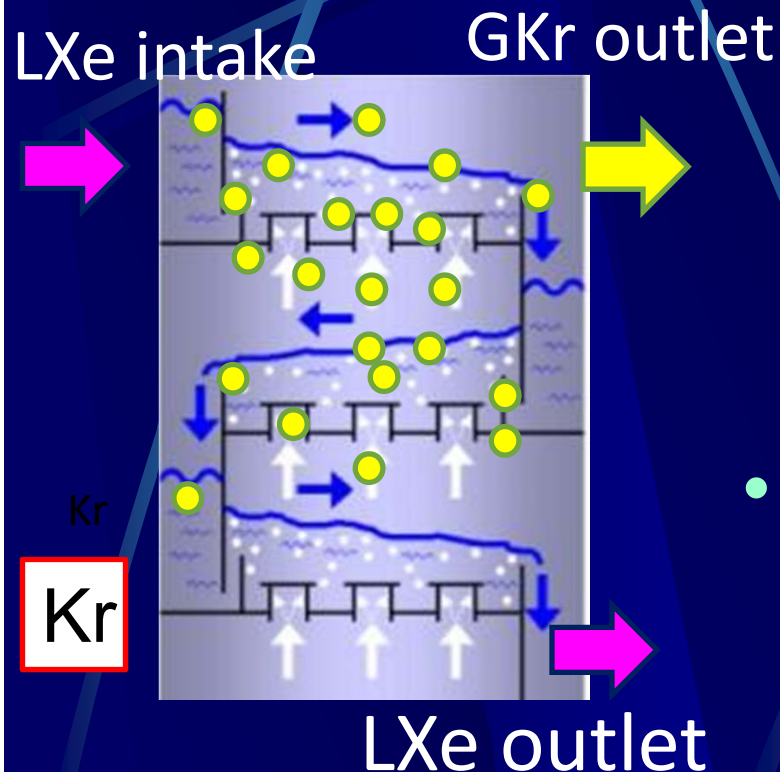
- Large Water Cherenkov OD -

- OD with 72 20"-PMTs \rightarrow active μ veto
- Large Vol. reduce γ and n



BG reduction(IV)

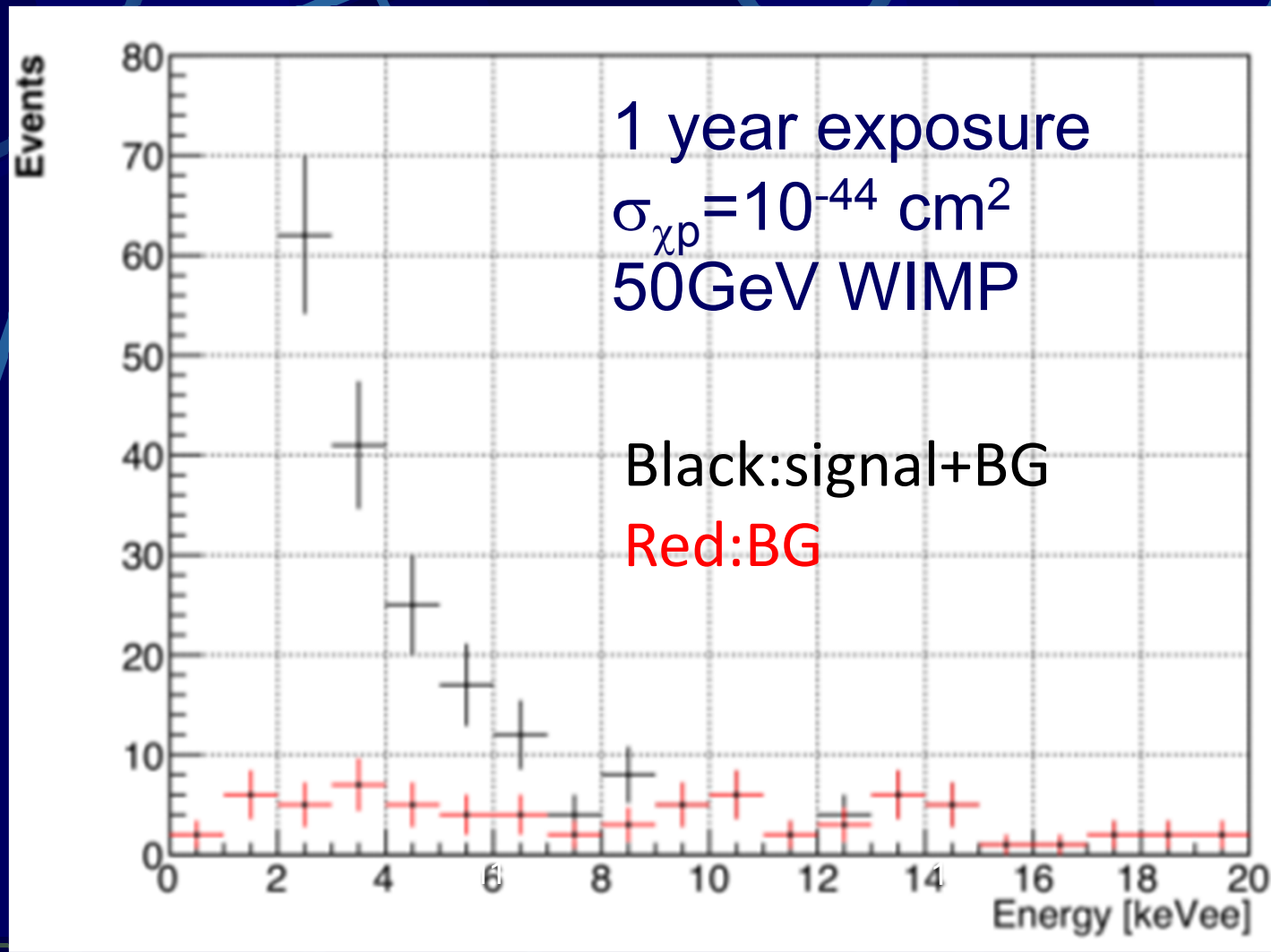
Xe purification



- Distillation:
 - 10^5 reduction for **Kr** (*)
 - **0.1ppm** \rightarrow **1ppt** for test sample
 - 1t of Xe was distilled at 10days before the filling.
- Other methods:
 - Getter
 - Filtering in liq. and gas phase is now under study for ^{222}Rn reduction.

(*) K. Abe et al. for XMASS collab., Astropart. Phys. 31 (2009) 290

Expected spectrum



Expected sensitivity

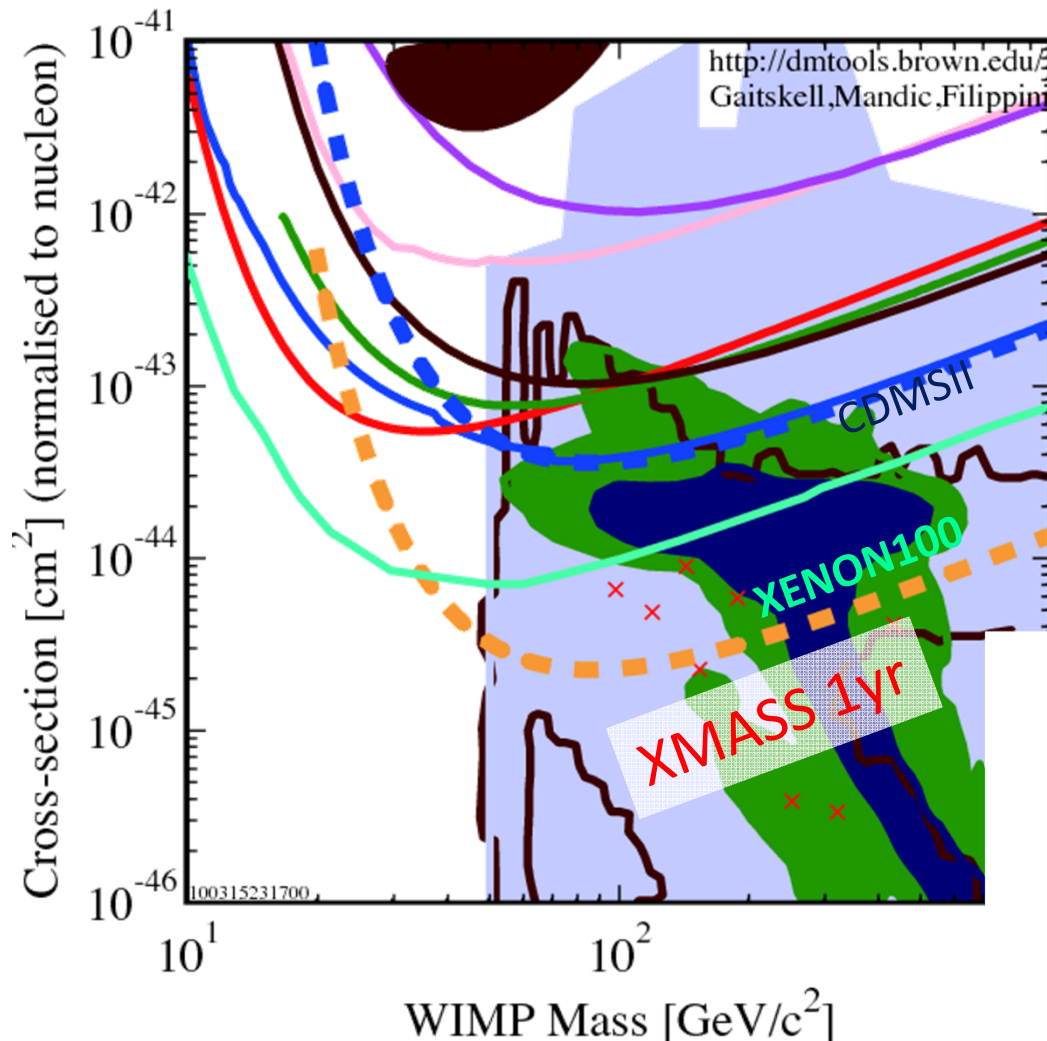
Spin Independent

$$\sigma_{cp} > 2 \times 10^{-45} \text{ cm}^2$$

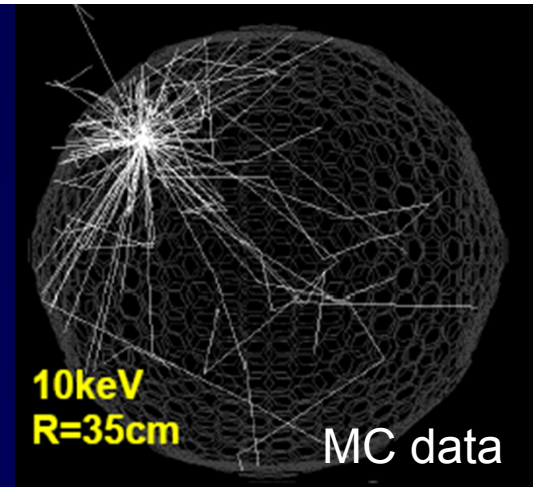
for 50-100 GeV WIMP
(90% C.L.)

Assumption:

- ✓ 1yr exposure,
- ✓ 100kg FV
- ✓ BG: 1×10^{-4} /keV/day/kg
- ✓ Q factor: 0.2



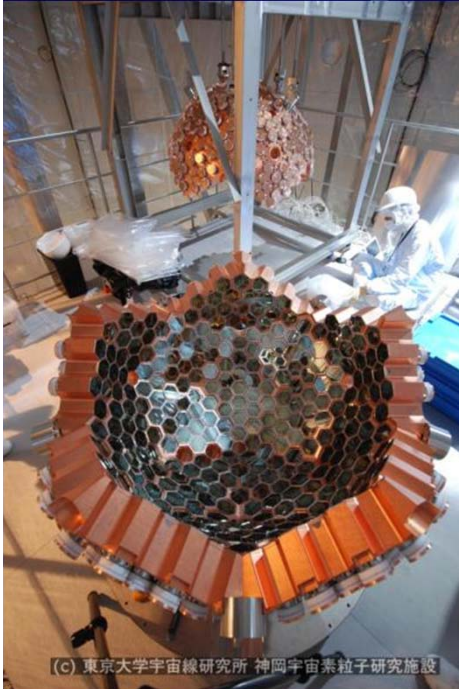
DATA listed top to bottom on plot
 DAMA/LIBRA 2008 3sigma, no ion channeling
 WARP 2.3L, 96.5 kg-days 55 keV threshold
 CRESST 2007 60 kg-day CaWO4
 Edelweiss II first result, 144 kg-days interleaved Ge
 ZEPLIN III (Dec 2008) result
 XENON10 2007, measured L_{eff} from Xe cube
 CDMS: Soudan 2004-2009 Ge
 Trotta et al 2008, CMSSM Bayesian: 68% contour
 Trotta et al 2008, CMSSM Bayesian: 95% contour
 Ellis et. al Theory region post-LEP benchmark points
 Baltz and Gondolo 2003
 Baltz and Gondolo, 2004, Markov Chain Monte Carlos
 100315231700



XMASS STATUS

- Performance check
- BG evaluation

Construction of XMASS

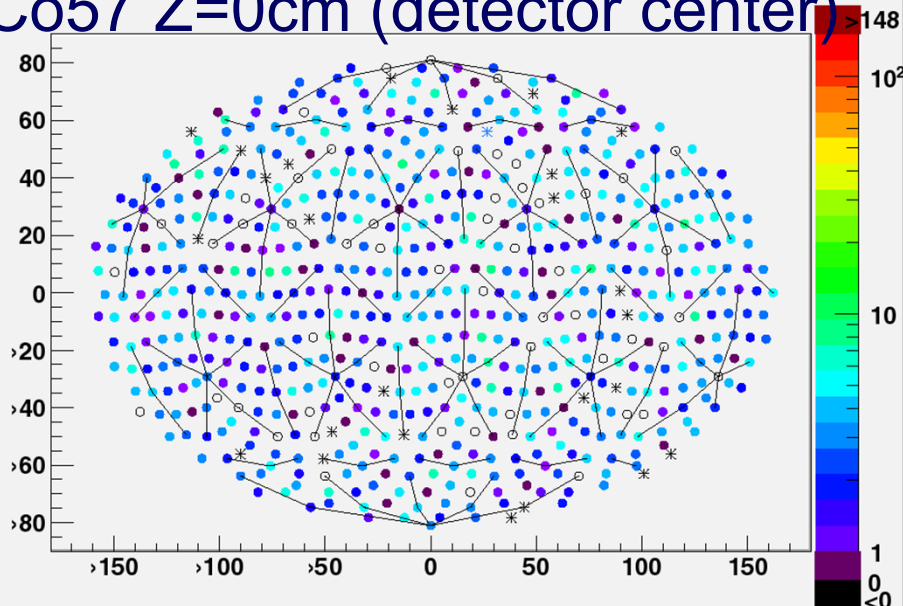


- 2010 Feb.: PMT installation was finished.
- 2010 Sep.: Detector assembly was finished.
- 2010 Sep.: Distillation and liq. xenon filling.

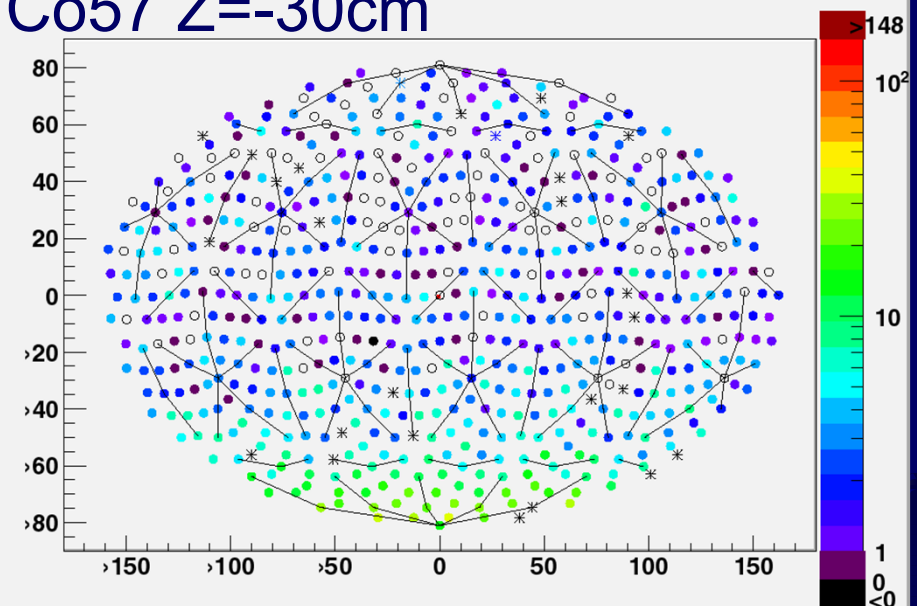
Current status of XMASS

- Everything was ready.
 - Detector itself
 - DAQ hardware and software
 - Calibration apparatus and so on
- 2010 Oct. → Commissioning phase

Co57 Z=0cm (detector center)



Co57 Z=-30cm



Calibration system

- Features
 - <1mm precision
 - Reducing “shadowing” by thin wire source
- Sources
 - ^{57}Co , ^{241}Am , ^{109}Cd , ^{55}Fe , ^{137}Cs , ...



Source rod with a dummy source

Stepping Motor

Linear Motion Feed-through

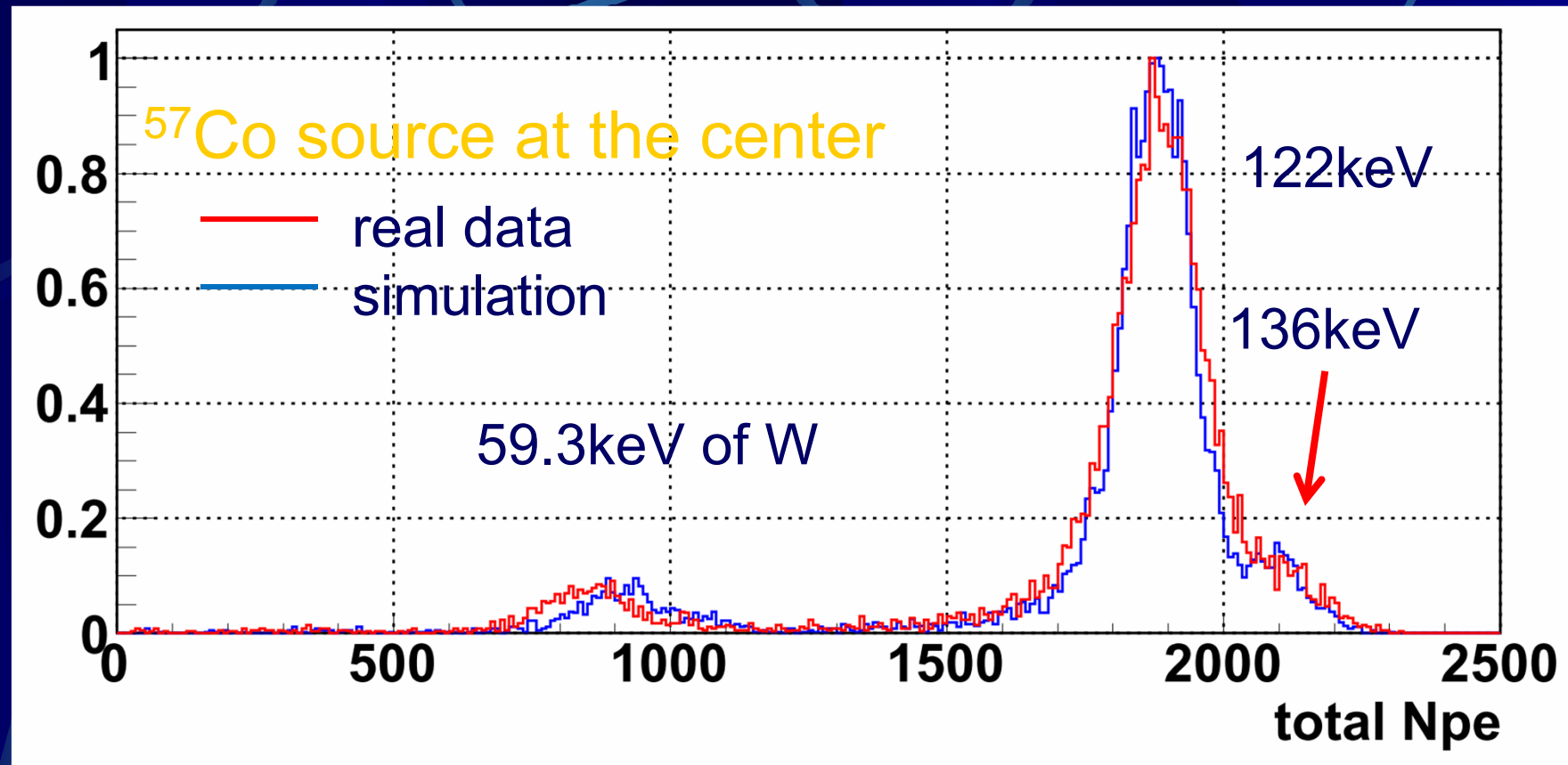
Source

Gate valve

~5m

Top
photo
tube

PE distribution

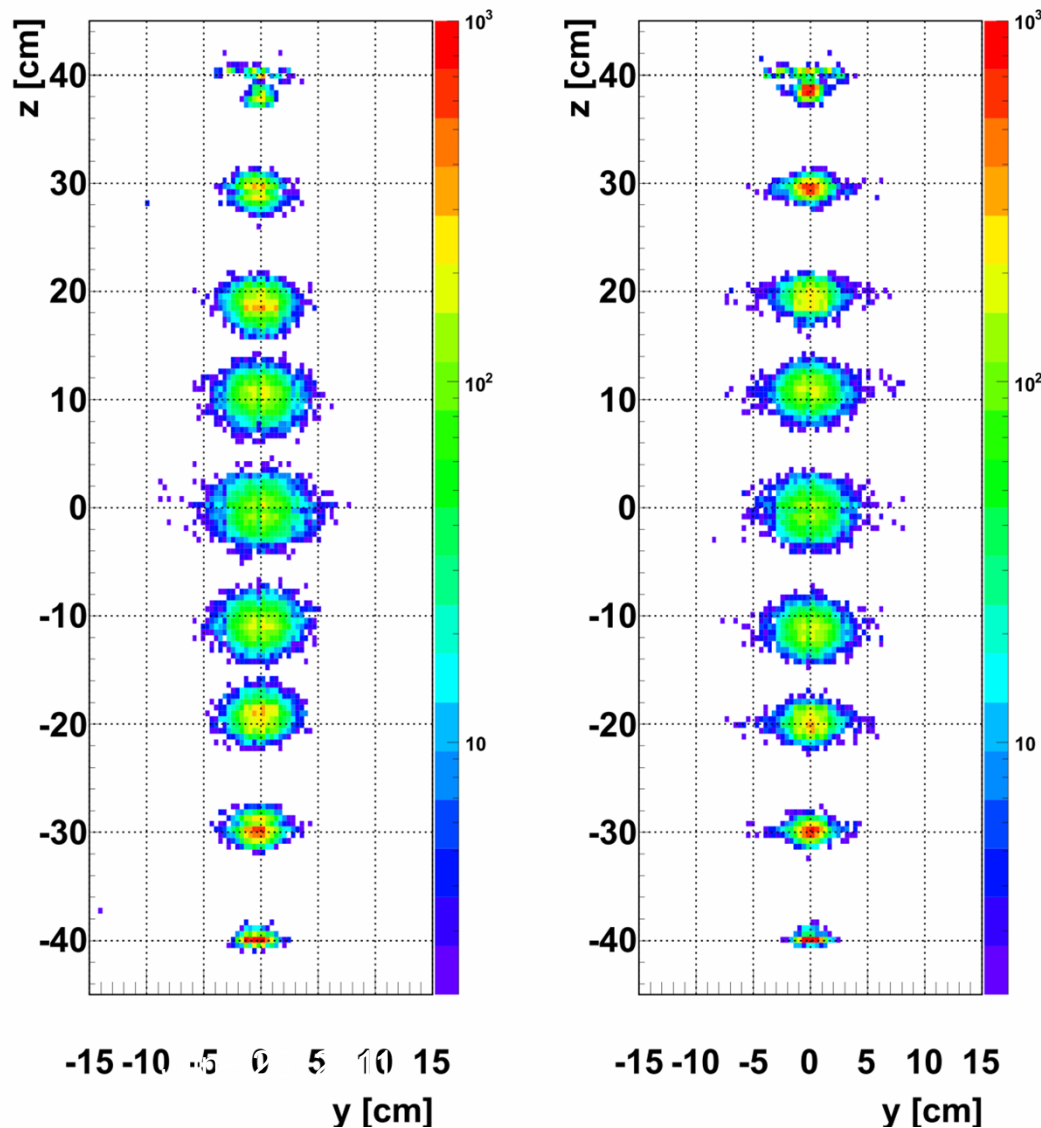


- High p.e. yield, $15.1 \pm 1.2 \text{ p.e./keV}$, was obtained.
- The photo electron yield distribution was well reproduced by MC.

Vertex reconstruction

Real Data

Simulation



r is well reconstructed.

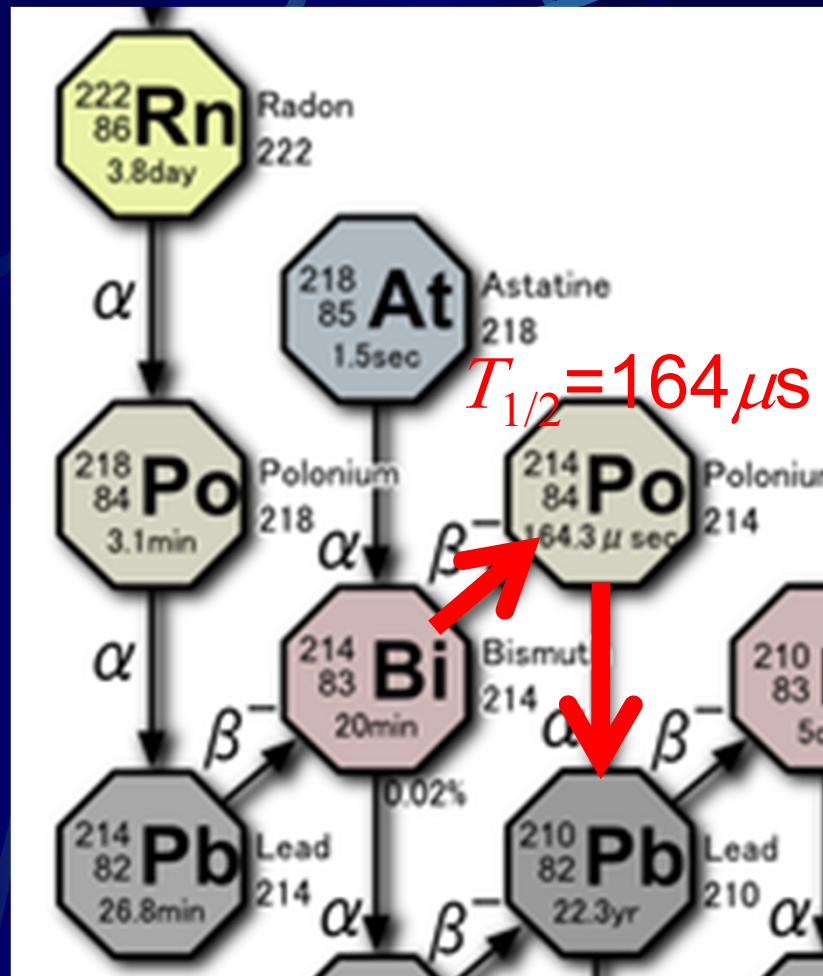
Δr is as expected by MC
1.4cm RMS @z=0cm
1.0cm RMS @+/-20cm

For 122keV γ rays

Evaluation of internal BG

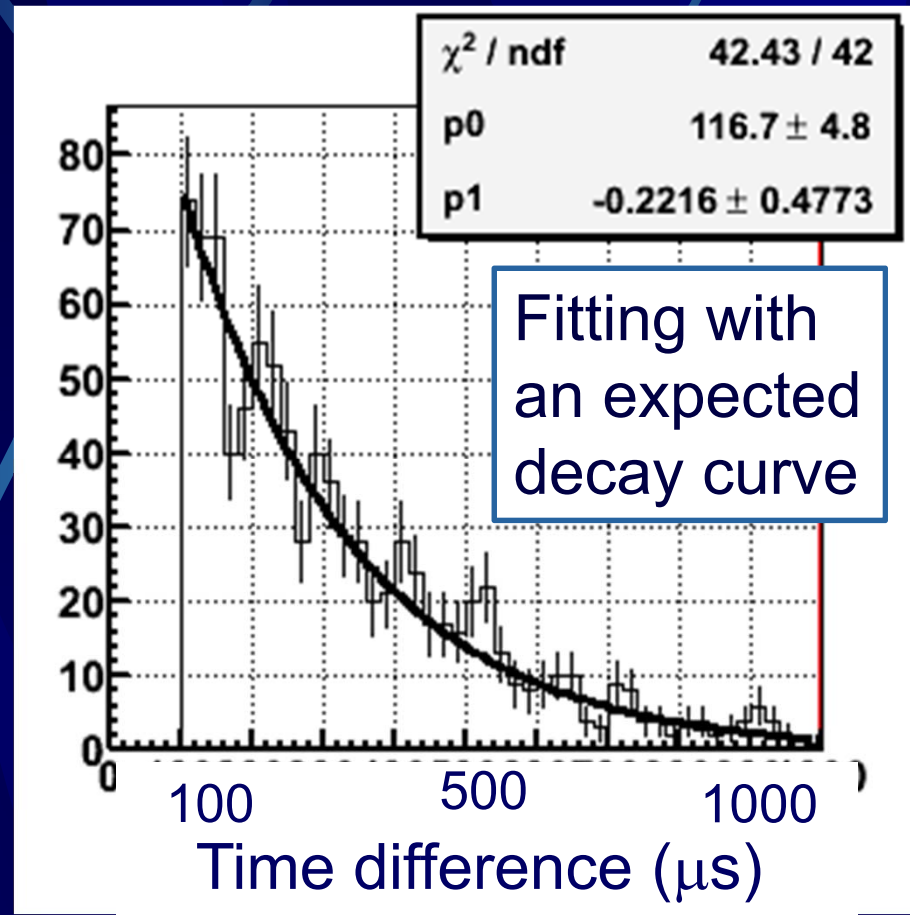
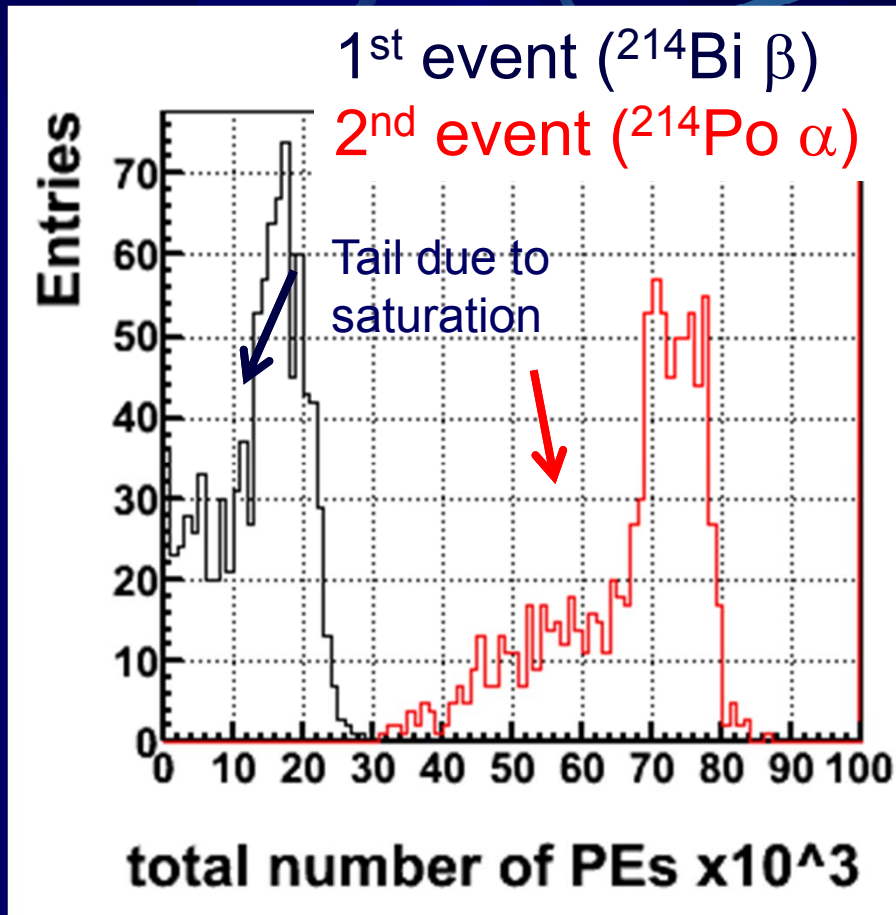
- External BG:
 - External BG (γ , n) can be effectively reduced by the water tank and the outer part of LXe.
- Internal BG:
 - Internal BG must be reduced by other means.
 - ^{222}Rn , ^{220}Rn , and ^{85}Kr are our concern because they give low energy BG.

^{222}Rn measurement



- Delayed coincidence
 - $T_{1/2} = 164 \mu\text{s}$
- The 2nd α is very bright, so gain of 321 PMTs are reduced for wide dynamic range.

^{222}Rn Result



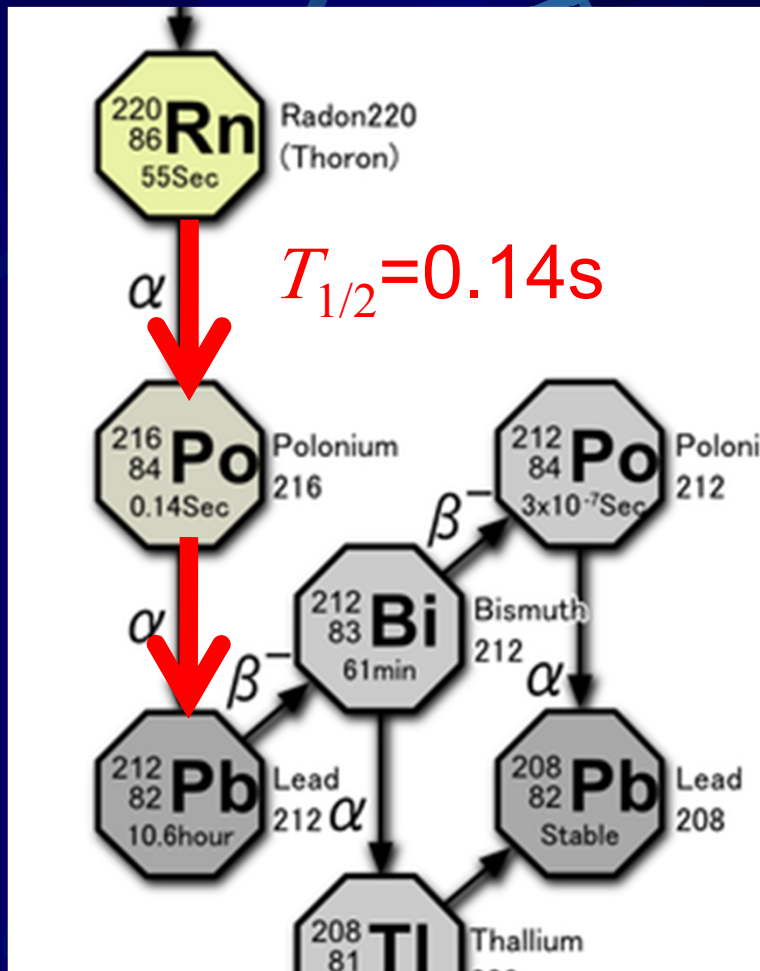
$^{222}\text{Rn}=8.2(0.5)\text{mBq}$ in inner volume

Target value=1mBq.

→ Investigating event distribution

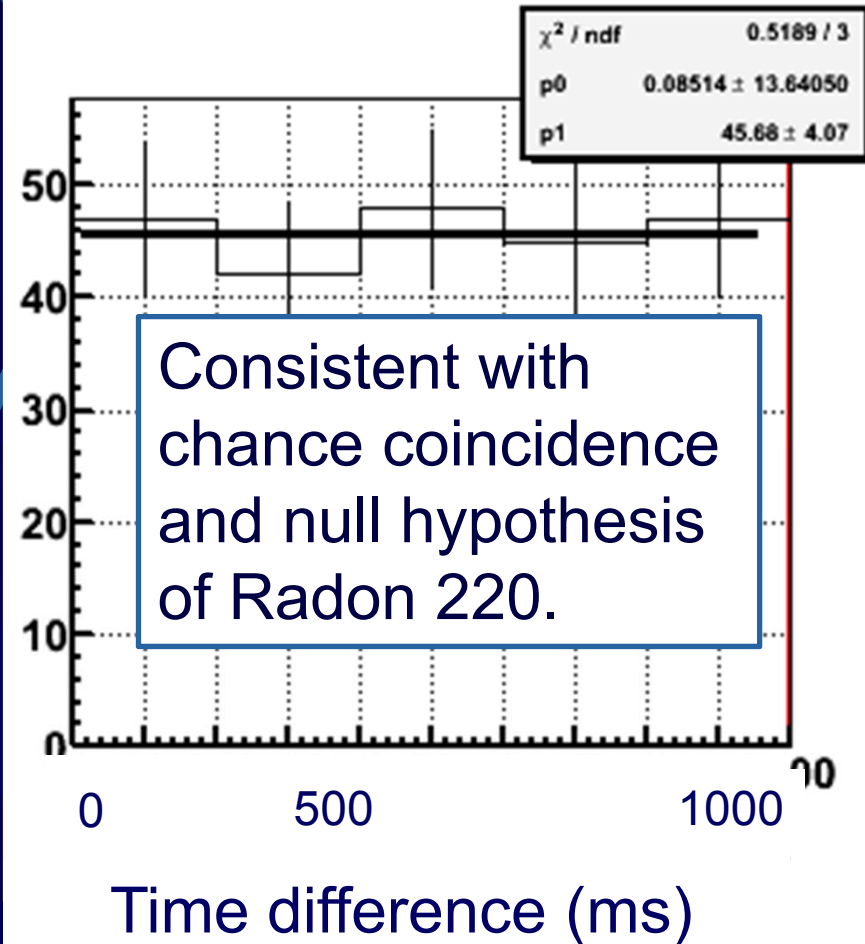
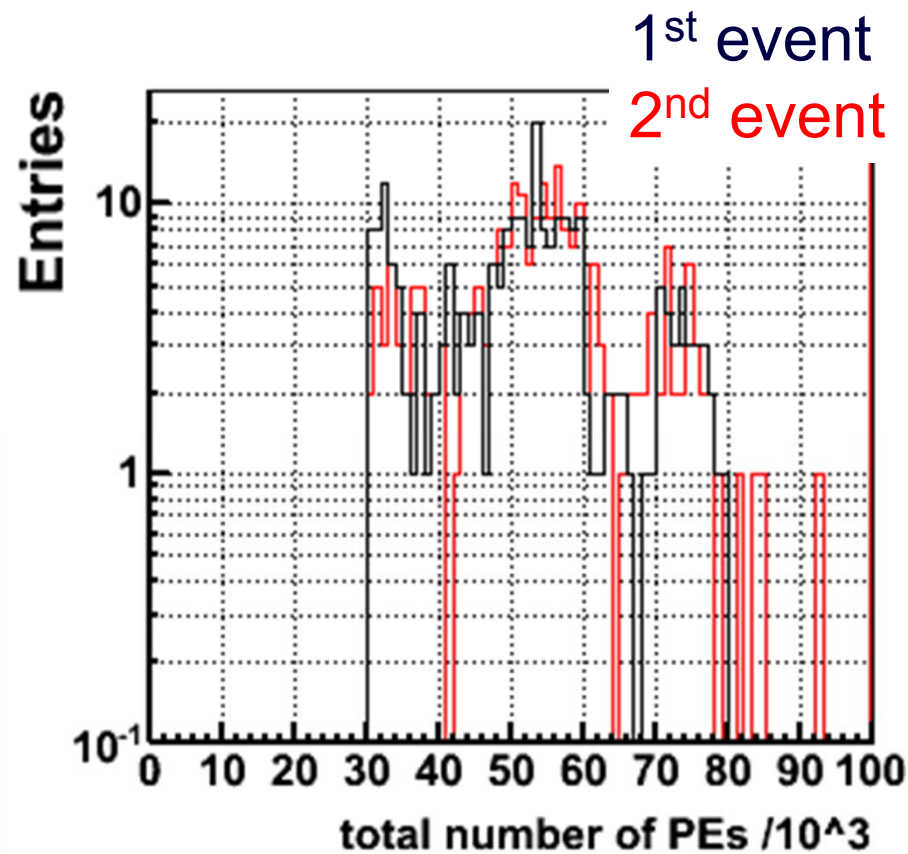
Considering measures to remove (Charcoal Trap)

^{220}Rn measurement



- Delayed coincidence
 - $T_{1/2}=0.14\text{s}$

^{220}Rn Result

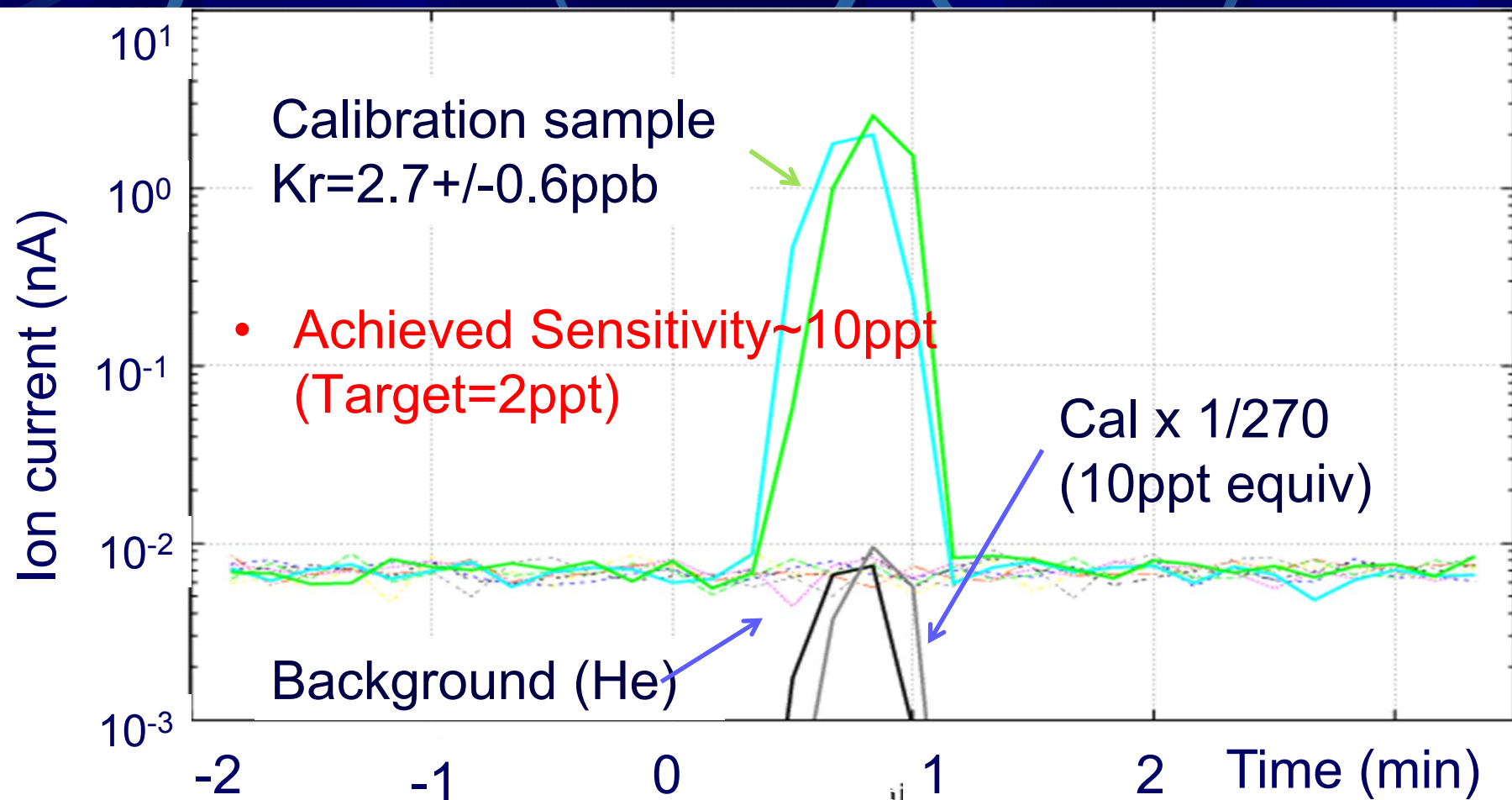


- $^{220}\text{Rn} < 0.28 \text{ mBq}$ in inner volume

Lower than our target value = 0.43 mBq .

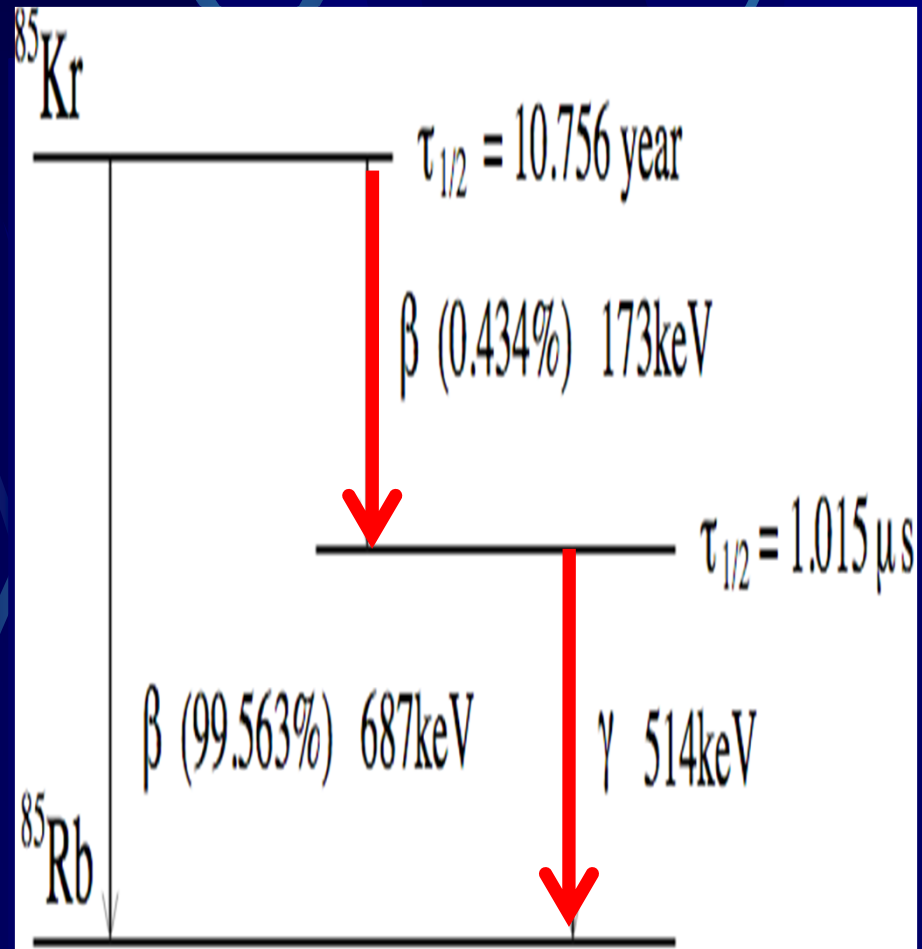
Kr measurement

- Gas Chromatography
+ Atmospheric Pressure Ionization Mass Spectrometer
- ^{85}Kr can be estimated with measured $^{85}\text{Kr}/\text{Kr}$



Kr measurement (II)

- Delayed coincident
- $T_{1/2} = 1.0 \mu s \rightarrow$
 ^{85}Kr measurement
with **FADC**
- New 650 FADCs in
November
 - They will boost our
detailed studies.



Summary

- The XMASS 800kg detector is for DM search with the sensitivity $2 \times 10^{-45} \text{cm}^2$ (spin independent case).
- Construction of the 800kg detector finished.
- Commissioning runs are on going to confirm performance and BG properties.
 - Energy resolution and vertex resolution were as expected.
 - $\sim 1 \text{cm}$ position resolution and $\sim 4\%$ energy resolution for $122 \text{keV } \gamma$.
 - Radon background are close to the target values
 - Kr contamination will be evaluated soon.
- Results are coming after commissioning.

BACKUP SLIDES

June 26, 2011

the 7th Patras workshop on
Axions, WIMPs and WISPs

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Vertex reconstruction

- $L(\mathbf{r}) = \prod_{i=\text{PMT}} \exp(-\mu(i)) \mu(i)^{n(i)} / n(i)!$
 - $\mu(i) = F(\mathbf{r}, i) / \sum_{i=\text{PMT}} F(\mathbf{r}, i) \times \text{total NPE}$
 - $n(i)$ = observed NPE in i -th PMT
 - $F(\mathbf{r}, i)$: Acceptance of i -th PMT
 - This is calculated by MC simulation.
- Find \mathbf{r} which gives min. $\log L(\mathbf{r})$

Energy reconstruction

- Calculated from
 - observed NPE
 - and total acceptance $\sum_{i=\text{PMT}} F(r, i)$

Why LXe?

- High Atomic mass Xe ($A \sim 131$) good for SI case ($\sigma \propto A^2$).
- Odd isotope (^{129}Xe (26.4%), ^{131}Xe (21.3%)) with large SD enhancement factors.
- High atomic number ($Z=54$) and density ($\rho \sim 3\text{g/cc}$)
 - -> Effective self-shielding.
 - -> Compact for large mass detector.
- High photon yield
- Easy to purify for both electro-negative and radioactive
 - impurities.
 - -> By circulation of Xe with getter for electro-negative.
 - -> Distillation for ^{85}Kr removal.

Reconstructed energy

^{57}Co source at the center

